

On Potentials Integrated by the Nikiforov–Uvarov Method

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ABSTRACT. We discuss basic potentials of the nonrelativistic and relativistic quantum mechanics that can be integrated in the Nikiforov and Uvarov paradigm with the aid of a computer algebra system. This approach may help the readers to study modern analytical methods of quantum physics.

Building on ideas of DE BROGLIE and EINSTEIN, I tried to show that the ordinary differential equations of mechanics, which attempt to define the co-ordinates of a mechanical system as functions of the time, are no longer applicable for “small” systems; instead there must be introduced a certain *partial* differential equation, which defines a variable ψ (“wave function”) as a function of the co-ordinates and the time.

Erwin Schrödinger [36]

1. Introduction

Discovery of the relativistic and nonrelativistic Schrödinger equations [33], [34], [35], [36], [37], [38] is discussed in [3] (see also the references therein). Invented about a century ago, the stationary Schrödinger equation turns out to have an enormously wide range of applications, from the quantum theory of atoms and molecules to solid state physics, quantum crystals, superfluidity, and superconductivity. Finding of the energy levels and the corresponding normalized wave functions of various systems is one of the basic problems of quantum physics. Only in a few elementary cases the exact solutions are known. They are usually investigated by different techniques. Nonetheless, those completely integrable problems are important in creation of mathematical models for complex quantum systems. Moreover, they may provide a useful testing ground for verification of numerical methods. A true story of calculations of the energy levels for the two-electron atoms [29] is presented in [20].

We assemble analytical solutions for a range of potentials in the nonrelativistic and relativistic quantum mechanics that are available in the literature. Data for most of the potentials that can be studied, in a unified way, by the so-called Nikiforov–Uvarov method [27] are collected, independently verified, and completed with the help of the Mathematica computer algebra system. Only bound states are discussed. On the contrary, in a traditional approach, for each of those problems one has to identify and factor out the singularities of the corresponding square integrable wave functions and find the remaining terminating power series expansions or use algebraic methods (see, for example, [4], [6], [7], [9], [10], [12], [17], [18],

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[23], [39], [40]). As a result, each of such problems has to be treated separately, which is not suitable for a unified computer algebra approach.

One should admit nonetheless that, despite of our successful unification and creation of the elaborate Mathematica codes, each of the collected cases, with exception of a few trivial ones, has required a separate consideration with lots of specific details that are far away from a formal application of the Nikiforov–Uvarov paradigm, as it might be thought of at the first glance. We hope that the interested reader, already confronted or about to, with all those quantization problems in the study, research or teaching will be able to appreciate our diligent work. Applications of hypergeometric functions are important to know for all mathematicians.

Moreover, an interesting new case of the modified Hulthén potential and its rotation correction can be useful in the study of vibration-rotation spectra of diatomic molecules; this is a work in progress. In addition, our analysis has revealed incomplete information and typos for some valuable potentials discussed in the classical sources. The complementary Mathematica notebook gives an opportunity to fix those typos immediately. In general, despite of a long history of applications of the computer algebra methods in quantum physics, we were unable to find suitable techniques in the available literature and, as a result, had to create most of our Mathematica codes from scratch, of course, by using our preliminary experience and enthusiasm in solving quantum physics and similar problems with the aid of computers.

Our review article is organized as follows. In the next section, we introduce the basics of the Nikiforov–Uvarov approach and then, successively, apply it to the main problems of introductory quantum mechanics, such as harmonic oscillators, Bessel functions, Coulomb problems, Pöschl–Teller potential holes, Kratzer’s molecular potential, Hulthén potentials, and Morse potentials, in the forthcoming sections. All calculations are verified in a complementary Mathematica notebook that can serve both educational and research purposes. Appendices A and B contain, for the reader’s convenience, the data for classical orthogonal polynomials and a useful integral evaluation, respectively, in order to make our presentation as self-contained as possible. Appendix C describes the Mathematica notebook.

This review is written for those who study quantum mechanics and would like to see more details than in the classical textbooks by utilizing the advanced computer algebra system, Mathematica. It is motivated by an introductory course in mathematics of quantum mechanics which one of the authors (SKS) has been teaching at Arizona State University for more than two decades.

2. Summary of the Nikiforov–Uvarov approach

The generalized equation of the hypergeometric type

$$(2.1) \quad u'' + \frac{\tilde{\tau}(x)}{\sigma(x)}u' + \frac{\tilde{\sigma}(x)}{\sigma^2(x)}u = 0$$

($\sigma(x)$, $\tilde{\sigma}(x)$ are polynomials of degrees at most 2 and $\tilde{\tau}(x)$ is a polynomial degree at most one) by the substitution

$$(2.2) \quad u = \varphi(x)y(x)$$

can be reduced to the form

$$(2.3) \quad \sigma(x)y'' + \tau(x)y' + \lambda y = 0$$

if:

$$(2.4) \quad \frac{\varphi'}{\varphi} = \frac{\pi(x)}{\sigma(x)}, \quad \pi(x) = \frac{1}{2}(\tau(x) - \tilde{\tau}(x))$$

(or, $\tau(x) = \tilde{\tau} + 2\pi$, for later),

$$(2.5) \quad k = \lambda - \pi'(x) \quad (\text{or, } \lambda = k + \pi'),$$

and

$$(2.6) \quad \pi(x) = \frac{\sigma' - \tilde{\tau}}{2} \pm \sqrt{\left(\frac{\sigma' - \tilde{\tau}}{2}\right)^2 - \tilde{\sigma} + k\sigma}$$

is a linear function. (Use the choice of the constant k to complete the square under the radical sign; see [27] and our argument below for more details.)

In Nikiforov–Uvarov’s method, the energy levels can be obtained from the quantization rule:

$$(2.7) \quad \lambda + n\tau' + \frac{1}{2}n(n-1)\sigma'' = 0 \quad (n = 0, 1, 2, \dots)$$

and the corresponding square-integrable solutions are classical orthogonal polynomials, up to a factor. They can be found by the *Rodrigues-type formula* [27]:

$$(2.8) \quad y_n(x) = \frac{B_n}{\rho(x)} (\sigma^n(x)\rho(x))^{(n)}, \quad (\sigma\rho)' = \tau\rho,$$

where B_n is a constant (see also [42] and Table 19). (The corresponding data for basic nonrelativistic and relativistic problems are presented in the Tables 1–18 below.)

Let us try to transform the differential equation (2.1) to the simplest form by the change of unknown function $u = \varphi(x)y$ with the help of some special choice of function $\varphi(x)$.

Substituting $u = \varphi(x)y$ in (2.1) one gets

$$(2.9) \quad y'' + \left(\frac{\tilde{\tau}}{\sigma} + 2\frac{\varphi'}{\varphi}\right)y' + \left(\frac{\tilde{\sigma}}{\sigma^2} + \frac{\tilde{\tau}\varphi'}{\sigma\varphi} + \frac{\varphi''}{\varphi}\right)y = 0.$$

Equation (2.9) should not be more complicated than our original equation (2.1). Thus, it is natural to assume that the coefficient in front of y' has the form $\tau(x)/\sigma(x)$, where $\tau(x)$ is a polynomial of degree at most one. This implies the following first-order differential equation

$$(2.10) \quad \frac{\varphi'}{\varphi} = \frac{\pi(x)}{\sigma(x)}$$

for the function $\varphi(x)$, where

$$(2.11) \quad \pi(x) = \frac{1}{2}(\tau(x) - \tilde{\tau}(x))$$

is a polynomial of degree at most one. As a result, equation (2.9) takes the form

$$(2.12) \quad y'' + \frac{\tau(x)}{\sigma(x)}y' + \frac{\bar{\sigma}(x)}{\sigma^2(x)}y = 0,$$

where

$$(2.13) \quad \bar{\sigma}(x) = \tilde{\sigma}(x) + \pi^2(x) + \pi(x)[\tilde{\tau}(x) - \sigma'(x)] + \pi'(x)\sigma(x).$$

The functions $\tau(x)$ and $\bar{\sigma}(x)$ are polynomials of degrees at most one and two in x , respectively. Therefore, equation (2.12) is an equation of the same type as our original equation (2.1).

By using a special choice of the polynomial $\pi(x)$ we can reduce (2.12) to the simplest form assuming that

$$(2.14) \quad \bar{\sigma}(x) = \lambda \sigma(x),$$

where λ is some constant. Then equation (2.12) takes the form (2.3). We call equation (2.3) a *differential equation of hypergeometric type* and its solutions *functions of hypergeometric type*. In this context, it is natural to call equation (2.1) a *generalized differential equation of hypergeometric type* [27].

The condition (2.14) can be rewritten as

$$(2.15) \quad \pi^2 + (\tilde{\tau} - \sigma') \pi + \tilde{\sigma} - k\sigma = 0,$$

where

$$(2.16) \quad k = \lambda - \pi'(z)$$

is a constant. Assuming that this constant is known, we can find $\pi(x)$ as a solution (2.6) of the quadratic equation (2.15). But $\pi(x)$ is a polynomial, therefore the second degree polynomial

$$(2.17) \quad p(x) = \left(\frac{\sigma'(x) - \tilde{\tau}(x)}{2} \right)^2 - \tilde{\sigma}(x) + k\sigma(x)$$

under the radical should be a square of a linear function and the discriminant of $p(x)$ should be zero. This condition gives an equation for the constant k , which is, generally, a quadratic equation. Given k as a solution of this equation, we find $\pi(x)$ by the quadratic formula (2.6), then $\tau(x)$ and λ by (2.11) and (2.16). Finally, we find the function $\varphi(x)$ as a solution of (2.10). It is clear that the reduction of equation (2.1) to the simplest form (2.3) can be accomplished by a few different ways in accordance with different choices of the constant k and different signs in (2.6) for $\pi(x)$.

A closed form for the constant k can be obtained as follows [3]. Let

$$(2.18) \quad p(x) = \left(\frac{\sigma' - \tilde{\tau}}{2} \right)^2 - \tilde{\sigma} + k\sigma = q(x) + k\sigma(x),$$

where

$$(2.19) \quad q(x) = \left(\frac{\sigma' - \tilde{\tau}}{2} \right)^2 - \tilde{\sigma}.$$

Completing the square, one gets

$$(2.20) \quad p(x) = \frac{p''}{2} \left(x + \frac{p'(0)}{p''} \right)^2 - \frac{(p'(0))^2 - 2p''p(0)}{2p''},$$

where the last term must be eliminated:

$$(2.21) \quad (p'(0))^2 - 2p''p(0) = 0.$$

Therefore,

$$(2.22) \quad (q'(0) + k\sigma'(0))^2 - 2(q'' + k\sigma'')(q(0) + k\sigma(0)) = 0,$$

which results in the following quadratic equation:

$$(2.23) \quad ak^2 + 2bk + c = 0.$$

Here,

$$(2.24) \quad a = (\sigma'(0))^2 - 2\sigma''\sigma(0),$$

$$(2.25) \quad b = q'(0)\sigma'(0) - \sigma''q(0) - \sigma(0)q'',$$

$$(2.26) \quad c = (q'(0))^2 - 2q''q(0).$$

Solutions are

$$(2.27) \quad k_0 = -\frac{c}{2b}, \quad \text{if } a = 0$$

and

$$(2.28) \quad k_{1,2} = \frac{-b \pm \sqrt{d}}{a}, \quad \text{if } a \neq 0.$$

Here,

$$(2.29) \quad d = b^2 - ac = (\sigma(0)q'' - \sigma''q(0))^2 - 2(\sigma'(0)q'' - \sigma''q'(0))(\sigma(0)q'(0) - \sigma'(0)q(0)).$$

As a result,

$$(2.30) \quad p(x) = \left[\frac{p''}{2} \left(x + \frac{p'(0)}{p''} \right)^2 \right]_{k=k_{0,1,2}},$$

which allows evaluating the linear function $\pi(x)$ in the Nikiforov–Uvarov technique.

Examples, for the most integrable cases that are available in the literature, are presented in the Tables 1–18 below, where all these analytical arguments are implemented in a supplementary Mathematica notebook (Appendix C).

“Quantizations” of some solutions of (2.3) require generalized power series expansions, given by

$$(2.31) \quad y(x) = (x-a)^\alpha \sum_{m=0}^{\infty} \prod_{k=0}^{m-1} \frac{(\lambda - \lambda_{\alpha+k})(a-x)}{\tau_{\alpha+k}(a)(\alpha+k+1)},$$

where

$$(2.32) \quad \tau_\mu(x) = \tau(x) + \mu\sigma'(x), \quad \lambda_\mu = -\mu\tau' - \mu(\mu-1)\sigma''/2,$$

provided that $\sigma(a) = 0$ and $\alpha\tau_{\alpha-1}(a) = 0$ (in particular when $\alpha = 0$). More details can be found in [42].

3. Harmonic Oscillator

Let us consider the one-dimensional stationary Schrödinger equation for the harmonic oscillator:

$$(3.1) \quad -\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} + \frac{1}{2}m\omega^2 x^2\psi = E\psi$$

with the orthonormal real-valued wave function

$$(3.2) \quad \int_{-\infty}^{\infty} \psi^2(x) dx = 1.$$

Introducing dimensionless variables

$$(3.3) \quad \psi(x) = u(\xi), \quad x = \xi \sqrt{\frac{\hbar}{m\omega}}, \quad E = \hbar\omega\varepsilon$$

one gets

$$(3.4) \quad u'' + (2\varepsilon - \xi^2) u = 0.$$

Here, $\sigma(\xi) = 1$, $\tilde{\tau}(\xi) = 0$, and $\tilde{\sigma}(\xi) = 2\varepsilon - \xi^2$. Therefore,

$$(3.5) \quad \pi(\xi) = \pm \sqrt{k - 2\varepsilon + \xi^2} = \pm \xi, \quad k = 2\varepsilon.$$

We pick $\pi = -\xi$, which gives a negative derivative for

$$(3.6) \quad \tau(\xi) = \tilde{\tau}(\xi) + 2\pi(\xi) = -2\xi.$$

Then

$$(3.7) \quad \frac{\varphi'}{\varphi} = \frac{\pi(\xi)}{\sigma(\xi)} = -\xi, \quad \varphi(\xi) = e^{-\xi^2/2}$$

and $\lambda = 2\varepsilon - 1$, $\rho(\xi) = e^{-\xi^2}$. The energy levels are $\varepsilon = \varepsilon_n = n + 1/2$, ($n = 0, 1, 2, \dots$) from (2.7). The eigenfunctions,

$$(3.8) \quad y_n(\xi) = B_n e^{\xi^2} \frac{d^n}{d\xi^n} \left(e^{-\xi^2} \right),$$

are, up to a normalization, the Hermite polynomials (Table 19).

As a result, the orthonormal wave functions are given by [34], [36], [23]

$$(3.9) \quad \psi(x) = \left(\frac{m\omega}{\pi\hbar} \right)^{1/4} \frac{1}{\sqrt{2^n n!}} \exp\left(-\frac{m\omega}{2\hbar} x^2\right) H_n \left(x \sqrt{\frac{m\omega}{\hbar}} \right),$$

corresponding to the discrete energy levels

$$(3.10) \quad E_n = \hbar\omega \left(n + \frac{1}{2} \right) \quad (n = 0, 1, 2, \dots),$$

in Gaussian units; see also Figure 1 for graphs of the first five wave functions. (More general, “missing”, solutions of the time-dependent Schrödinger equation are discussed in [19], [21], and [24].)

4. Bessel Functions

Let us also mention some solutions of the *Bessel equation*:

$$(4.1) \quad z^2 u'' + z u' + (z^2 - \nu^2) u = 0.$$

With the aid of the change of the function $u = \varphi(z)y$ when $\varphi(z) = z^\nu e^{iz}$ this equation can be reduced to the hypergeometric form

$$(4.2) \quad z y'' + (2iz + 2\nu + 1) y' + i(2\nu + 1) y = 0$$

(Table 2) and one can obtain the *Poisson integral representations* for the Bessel functions of the first kind, $J_\nu(z)$, and the Hankel functions of the first and second

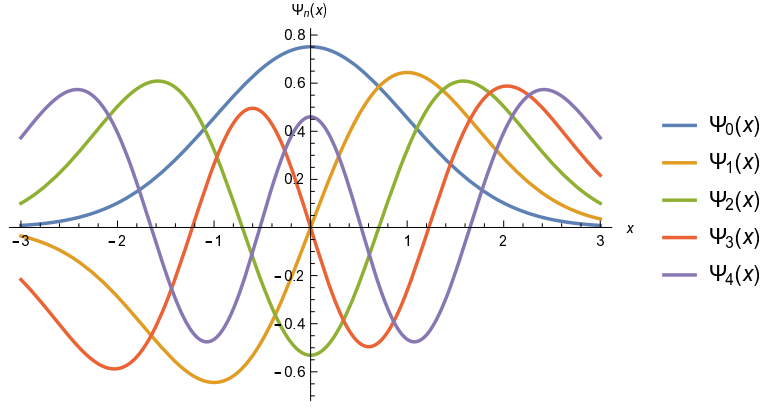


FIGURE 1. “The first five proper vibrations of the Planck oscillator according to undulatory mechanics. Outside the region $-3 \leq x \leq 3$, represented here, all five functions approach x -axis in monotonic fashion.”— generated by Mathematica following the original Schrödinger article [36].

TABLE 1. Stationary Schrödinger equation for the harmonic potential $U(x) = \frac{1}{2}m\omega^2x^2$.

$\sigma(\xi)$	1
$\tilde{\sigma}(\xi)$	$2\varepsilon - \xi^2$
$\tilde{\tau}(\xi)$	0
k	2ε
$\pi(\xi)$	$\pm\xi$
$\tau(\xi) = \tilde{\tau} + 2\pi$	-2ξ
$\lambda = k + \pi'$	$2\varepsilon - 1$
$\varphi(\xi)$	$e^{-\xi^2/2}$
$\rho(\xi)$	$e^{-\xi^2}$
$y_n(\xi)$	$C_n H_n(\xi)$
C_n^2	$\frac{1}{\sqrt{\pi} 2^n n!}$

kind, $H_\nu^{(1)}(z)$ and $H_\nu^{(2)}(z)$:

$$(4.3) \quad J_\nu(z) = \frac{(z/2)^\nu}{\sqrt{\pi} \Gamma(\nu + 1/2)} \int_{-1}^1 (1-t^2)^{\nu-1/2} \cos(zt) dt,$$

$$(4.4) \quad H_\nu^{(1,2)}(z) = \sqrt{\frac{2}{\pi z}} \frac{e^{\pm i(z - \frac{\pi}{2}\nu - \frac{\pi}{4})}}{\Gamma(\nu + 1/2)} \int_0^\infty e^{-t} t^{\nu-1/2} \left(1 \pm \frac{it}{2z}\right)^{\nu-1/2} dt,$$

TABLE 2. Bessel's equation.

$\sigma(z)$	z
$\tilde{\sigma}(z)$	$z^2 - \nu^2$
$\tilde{\tau}(z)$	1
k	$\pm 2i\nu$
$\pi(z)$	$\pm \nu \pm iz$
$\tau(z) = \tilde{\tau} + 2\pi$	$1 + 2\nu + 2iz$
$\lambda = k + \pi'$	$i(2\nu + 1)$
$\varphi(z)$	$z^{\pm \nu} e^{\pm iz}$
$\rho(z)$	$z^{2\nu} e^{2iz}$

where $\operatorname{Re} \nu > -1/2$. It is then possible to deduce from these integral representations all the remaining properties of these functions. (For details, see [27] and [42] or [44] and [45]).

5. Central Field: Spherical Harmonics

The stationary Schrödinger equation in the central field with the potential energy $U(r)$ is given by

$$(5.1) \quad \Delta \psi + \frac{2m}{\hbar^2} (E - U(r)) \psi = 0.$$

The Laplace operator in the spherical coordinates r, θ, φ has the form [26], [27]:

$$(5.2) \quad \Delta = \Delta_r + \frac{1}{r^2} \Delta_\omega$$

with

$$(5.3) \quad \Delta_r = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \right), \quad \Delta_\omega = \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial \varphi^2}.$$

and separation of the variables $\psi = R(r)Y(\theta, \varphi)$ gives

$$(5.4) \quad \Delta_\omega Y(\theta, \varphi) + \mu Y(\theta, \varphi) = 0,$$

$$(5.5) \quad \frac{1}{r^2} \frac{d}{dr} \left(r^2 \frac{dR(r)}{dr} \right) + \left(\frac{2m}{\hbar^2} (E - U(r)) - \frac{\mu}{r^2} \right) R(r) = 0.$$

Bounded single-valued solutions of equation (5.4) on the sphere S^2 exist only when $\mu = l(l+1)$ with $l = 0, 1, 2, \dots$. They are the spherical harmonics $Y = Y_{lm}(\theta, \varphi)$.

Looking for solutions in the form $Y = e^{im\varphi} f(\theta)$ with $m = 0, \pm 1, \pm 2, \dots$ one gets

$$(5.6) \quad \frac{1}{\sin \theta} \frac{d}{d\theta} \left(\sin \theta \frac{dg}{d\theta} \right) - \frac{m^2}{\sin^2 \theta} g + \mu g = 0.$$

The following change of variables $\xi = \cos \theta$ and $F(\xi) = f(\theta)$ results in the generalized equation of hypergeometric type:

$$(5.7) \quad (1 - \xi^2) F'' - 2\xi F' + \left(\mu - \frac{m^2}{1 - \xi^2} \right) F = 0,$$

with $\sigma(\xi) = 1 - \xi^2$, $\tilde{\tau}(\xi) = -2\xi$ and $\tilde{\sigma}(\xi) = \mu(1 - \xi^2) - m^2$, which can be reduced to the simpler form by the standard substitution $F = \varphi y$:

$$(5.8) \quad (1 - \xi^2) y'' - 2(|m| + 1)\xi y' + (\mu - |m|(|m| + 1))y = 0$$

Indeed, by (2.6)

$$(5.9) \quad \pi(\xi) = \pm \sqrt{(\mu - k)\xi^2 + k + m^2 - \mu},$$

or

$$(5.10) \quad \pi(\xi) = \begin{cases} \pm|m|, & k = \mu \\ \pm|m|\xi, & k = \mu - m^2 \end{cases}$$

where we should choose the case when the linear function $\tau = \tilde{\tau} + 2\pi$ will have a negative derivative and a zero on the interval $(-1, 1)$.

Then

$$(5.11) \quad \frac{\varphi'}{\varphi} = \frac{-|m|\xi}{1 - \xi^2}, \quad \ln \varphi = -|m| \int \frac{\xi d\xi}{1 - \xi^2} = \frac{1}{2} \ln(1 - \xi^2)$$

and

$$(5.12) \quad \varphi(\xi) = (1 - \xi^2)^{|m|/2} = (\sin \theta)^{|m|}.$$

(see our complementary Mathematica notebook and Table 3 for further details of calculations).

The final result is given by

$$(5.13) \quad Y_{lm}(\theta, \varphi) = A_m \frac{e^{im\varphi}}{2^{|m|} l!} \sqrt{\frac{2l+1}{4\pi}} (l-m)!(l+m)! (\sin \theta)^{|m|} P_{l-|m|}^{(|m|, |m|)}(\cos \theta).$$

Here, $P_n^{(\alpha, \alpha)}(\xi)$ are the Jacobi polynomials (Table 19); $A_m = (-1)^m$, $m \geq 0$ and $A_m = 1$, $m < 0$, or $A_m = (-1)^{(m+|m|)/2}$. (See [26], [27], [43] for more details.)

Solutions of the Laplace equation, $\Delta u = 0$, in the spherical coordinates, have the form

$$(5.14) \quad u = C_1 r^l Y_{lm}(\theta, \varphi) + C_2 r^{-l-1} Y_{lm}(\theta, \varphi),$$

where C_1 and C_2 are constants. Here, $r^l Y_{lm}(\theta, \varphi)$ are homogeneous polynomials in x, y, z [43]:

$$(5.15) \quad r^l Y_{lm}(\theta, \varphi) = \sqrt{\frac{2l+1}{4\pi}} (l+m)!(l-m)! \sum_{p,q,r} \frac{1}{p!q!r!} \left(\frac{x+iy}{2}\right)^p \left(\frac{x-iy}{2}\right)^q z^r,$$

where p, q, r are positive integers subject to the following conditions $p + q + r = l$ and $p - q = m$. (See our Mathematica notebook for direct verifications.)

6. Nonrelativistic Coulomb Problem

In view of identity

$$(6.1) \quad \frac{1}{r^2} \frac{d}{dr} \left(r^2 \frac{dR}{dr} \right) = \frac{1}{r} \frac{d^2}{dr^2} (rR),$$

the substitution $F(r) = rR(r)$ into (5.5) results in the standard radial equation

$$(6.2) \quad F'' + \left[\frac{2m_e}{\hbar^2} (E - U(r)) - \frac{l(l+1)}{r^2} \right] F = 0, \quad U(r) = -\frac{Ze^2}{r}$$

TABLE 3. Equation for spherical harmonics.

$\sigma(\xi)$	$1 - \xi^2, \quad \xi = \cos \theta, \quad 0 \leq \theta \leq \pi$
$\tilde{\sigma}(\xi)$	$\mu(1 - \xi^2) - m^2$
$\tilde{\tau}(\xi)$	-2ξ
k	$\mu - m^2$
$\pi(\xi)$	$- m \xi$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$-2(m + 1)\xi$
$\lambda = k + \pi'$	$\mu - m (m + 1)$
$\varphi(\xi)$	$(1 - \xi^2)^{ m /2} = (\sin \theta)^{ m }$
$\rho(\xi)$	$(1 - \xi^2)^{ m }$
$y_n(\xi)$	$N_{lm} P_{l- m }^{(m , m)}(\cos \theta), \quad n = l - m $
N_{lm}	$\frac{1}{2^{ m } l!} \sqrt{\frac{2l+1}{2} (l-m)!(l+m)!}$

for the nonrelativistic Coulomb problem in spherical coordinates. In dimensionless units,

$$(6.3) \quad F(r) = u(x), \quad x = \frac{r}{a_0}, \quad \varepsilon_0 = \frac{E}{E_0} \quad \left(a_0 = \frac{\hbar^2}{m_e e^2}, \quad E_0 = \frac{e^2}{a_0} \right)$$

the radial equation is a generalized equation of hypergeometric type,

$$(6.4) \quad u'' + \left[2 \left(\varepsilon_0 + \frac{Z}{x} \right) - \frac{l(l+1)}{x^2} \right] u = 0,$$

where

$$(6.5) \quad \sigma(x) = x, \quad \tilde{\tau}(x) = 0, \quad \tilde{\sigma}(x) = 2\varepsilon_0 x^2 + 2Zx - l(l+1).$$

Therefore, one can utilize Nikiforov and Uvarov's approach in order to determine the corresponding wave functions and discrete energy levels.

We transform (6.4) to the equation of hypergeometric type (2.3). The linear function $\pi(x)$ takes the form

$$(6.6) \quad \pi(x) = \frac{1}{2} \pm \sqrt{\frac{1}{4} - 2\varepsilon_0 x^2 - 2x + l(l+1) + kx},$$

or

$$(6.7) \quad \pi(x) = \frac{1}{2} \pm \begin{cases} \sqrt{-2\varepsilon_0} x + l + 1/2, & k = 2Z + (2l+1)\sqrt{-2\varepsilon_0} \\ \sqrt{-2\varepsilon_0} x - l - 1/2, & k = 2Z - (2l+1)\sqrt{-2\varepsilon_0} \end{cases}$$

where we should choose the case when the linear function $\tau = \tilde{\tau} + 2\pi$ will have a negative derivative and a zero on $(0, +\infty)$:

$$\tau(x) = 2(l+1 - x\sqrt{-2\varepsilon_0}).$$

This choice corresponds to

$$\pi(x) = l+1 - x\sqrt{-2\varepsilon_0}, \quad \varphi(x) = x^{l+1} \exp(-x\sqrt{-2\varepsilon_0})$$

and

$$\lambda = k + \pi' = 2 [Z - (l + 1)\sqrt{-2\varepsilon_0}].$$

The energy values are given by (2.7):

$$(6.8) \quad \varepsilon_0 = \frac{E}{E_0} = -\frac{Z^2}{2(n_r + l + 1)^2}, \quad E_0 = \frac{e^2}{a_0}.$$

Here, $n = n_r + l + 1$ is known as the principal quantum number.

In order to use the Rodrigues formula, one finds

$$\frac{\rho'}{\rho} = \frac{\tau - \sigma'}{\sigma} = \frac{2l + 1}{x} - \frac{2Z}{n},$$

or

$$\rho(x) = x^{2l+1} \exp\left(-\frac{2Z}{n}x\right), \quad x = \frac{r}{a_0}.$$

Therefore,

$$(6.9) \quad y_{n_r}(x) = \frac{B_{n_r}}{x^{2l+1}e^{-\eta}} \frac{d^{n_r}}{dx^{n_r}} (x^{n_r+2l+1}e^{-\eta}) = L_{n_r}^{2l+1}(\eta),$$

where

$$\eta = \frac{2Z}{n}x = \frac{2Z}{n} \left(\frac{r}{a_0}\right) = 2x\sqrt{-2\varepsilon_0},$$

and, up to a constant,

$$(6.10) \quad F(r) = rR(r) = C_{nl} \eta^{l+1} e^{-\eta/2} L_{n_r}^{2l+1}(\eta).$$

In view of the normalization condition

$$1 = \int_0^\infty F^2 dr = C_{nl}^2 \left(\frac{na_0}{2Z}\right) \int_0^\infty \eta^{2l+2} e^{-\eta} [L_{n_r}^{2l+1}(\eta)]^2 d\eta,$$

the three-term recurrence relation

$$\eta L_n^\alpha = -(n+1)L_{n+1}^\alpha + (\alpha + 2n + 1)L_n^\alpha - (\alpha + n)L_{n-1}^\alpha,$$

and the orthogonality property of the Laguerre polynomials (Table 19), one gets

$$(6.11) \quad C_{nl}^2 = \frac{Z}{a_0 n^2} \frac{(n-l-1)!}{(n+l)!}.$$

More details can be found in [27] and [42]. (See also Appendices A and B.)

As a result, the nonrelativistic Coulomb wave functions obtained by the method of separation of the variables in spherical coordinates, see above, are

$$(6.12) \quad \psi = \psi_{nlm}(\mathbf{r}) = R_{nl}(r) Y_{lm}(\theta, \varphi),$$

where $Y_{lm}(\theta, \varphi)$ are the spherical harmonics, the radial functions $R_{nl}(r)$ are given in terms of the Laguerre polynomials (Table 19) [4], [10], [23], [27], [39]:

$$(6.13) \quad R(r) = R_{nl}(r) = \frac{2}{n^2} \left(\frac{Z}{a_0}\right)^{3/2} \sqrt{\frac{(n-l-1)!}{(n+l)!}} e^{-\eta/2} \eta^l L_{n-l-1}^{2l+1}(\eta)$$

with

$$(6.14) \quad \eta = \frac{2Z}{n} \left(\frac{r}{a_0}\right), \quad a_0 = \frac{\hbar^2}{m_e e^2}$$

and the normalization is

$$(6.15) \quad \int_0^\infty R_{nl}^2(r) r^2 dr = 1.$$

TABLE 4. The Schrödinger equation for the Coulomb potential $U(r) = -Ze^2/r$. Dimensionless quantities: $r = a_0x$, $a_0 = \hbar^2/(m_e e^2) \simeq 0.5 \cdot 10^{-8}$ cm, $E_0 = e^2/a_0$, $R(r) = F(x) = u(x)/x$.

$\sigma(x)$	x
$\tilde{\sigma}(x)$	$2\varepsilon_0 x^2 + 2Zx - l(l+1)$, $\varepsilon_0 = E/E_0$
$\tilde{\tau}(x)$	0
k	$2Z - (2l+1)\sqrt{-2\varepsilon_0}$
$\pi(x)$	$l+1 - \sqrt{-2\varepsilon_0}x$
$\tau(x) = \tilde{\tau} + 2\pi$	$2(l+1 - \sqrt{-2\varepsilon_0}x)$
$\lambda = k + \pi'$	$2(Z - (l+1)\sqrt{-2\varepsilon_0})$
$\varphi(x)$	$x^{l+1}e^{-x\sqrt{-2\varepsilon_0}}$
$\rho(x)$	$x^{2l+1}e^{-(2Zx)/n}$, $x = r/a_0$
$y_{n_r}(x)$	$C_{n_r} L_{n_r}^{2l+1}(\eta)$, $\eta = \frac{2Z}{n}x = \frac{2Z}{n}\left(\frac{r}{a_0}\right)$
$C_{n_r}^2$	$\frac{Z}{a_0 n^2} \frac{(n-l-1)!}{(n+l)!}$, $n_r = n-l-1$

Here $n = 1, 2, 3, \dots$ is the principal quantum number of the hydrogen-like atom in the nonrelativistic Schrödinger theory; $l = 0, 1, \dots, n-1$ and $m = -l, -l+1, \dots, l-1, l$ are the quantum numbers of the angular momentum and its projection on the z -axis, respectively. The corresponding discrete energy levels in the cgs units are given by Bohr's formula [33]:

$$(6.16) \quad E = E_n = -\frac{m_e Z^2 e^4}{2\hbar^2 n^2},$$

where $n = 1, 2, 3, \dots$ is the principal quantum number; they do not depend on the quantum number of the orbital angular momenta l .

Remark. In the original Nikiforov–Uvarov approach, the variable coefficients $\sigma(x)$ and $\tau(x)$ in (2.3) should not depend on the eigenvalue λ . Here, we obtain

$$(6.17) \quad xy'' + 2(l+1 - x\sqrt{-2\varepsilon_0})y' + 2(Z - (l+1)\sqrt{-2\varepsilon_0})y = 0.$$

Nonetheless, the change of variables $y(x) = Y(\eta)$ with $\eta = 2x\sqrt{-2\varepsilon_0}$ results in

$$(6.18) \quad \eta Y'' + (2l+2 - \eta)Y' + \left(\frac{Z}{\sqrt{-2\varepsilon_0}} - l - 1\right)Y = 0.$$

Thus the Nikiforov–Uvarov method can be applied and the uniqueness of square integrable solutions holds.

In a similar fashion, one can consider solution of the Kepler problem in the so-called parabolic coordinates, which is important in the theory of Stark effect [35], [23].

7. Relativistic Schrödinger Equation

The stationary relativistic Schrödinger equation has the form [3], [7], [17], [39]:

$$(7.1) \quad \left(E + \frac{Ze^2}{r}\right)^2 \chi = (-\hbar^2 c^2 \Delta + m^2 c^4) \chi.$$

We separate the variables in spherical coordinates, $\chi(r, \theta, \varphi) = R(r)Y_{lm}(\theta, \varphi)$, where $Y_{lm}(\theta, \varphi)$ are the spherical harmonics with familiar properties [43]. As a result,

$$(7.2) \quad \frac{1}{r^2} \frac{d}{dr} \left(r^2 \frac{dR}{dr} \right) + \left[\frac{(E + Ze^2/r)^2 - m^2 c^4}{\hbar^2 c^2} - \frac{l(l+1)}{r^2} \right] R = 0 \quad (l = 0, 1, 2, \dots).$$

In the dimensionless quantities,

$$(7.3) \quad \varepsilon = \frac{E}{mc^2}, \quad x = \beta r = \frac{mc}{\hbar} r, \quad \mu = \frac{Ze^2}{\hbar c},$$

for the new radial function,

$$(7.4) \quad R(r) = F(x) = \frac{u(x)}{x},$$

one gets

$$(7.5) \quad \frac{1}{x^2} \frac{d}{dx} \left(x^2 \frac{dF}{dx} \right) + \left[\left(\varepsilon + \frac{\mu}{x} \right)^2 - 1 - \frac{l(l+1)}{x^2} \right] F = 0.$$

Given the identity $(x^2 F')' = x(xF)''$, we obtain

$$(7.6) \quad u'' + \left[\left(\varepsilon + \frac{\mu}{x} \right)^2 - 1 - \frac{l(l+1)}{x^2} \right] u = 0.$$

This is the generalized equation of hypergeometric form (2.1), when

$$(7.7) \quad \sigma(x) = x, \quad \tilde{\tau}(x) \equiv 0, \quad \tilde{\sigma}(x) = (\varepsilon^2 - 1)x^2 + 2\mu\varepsilon x + \mu^2 - l(l+1)$$

The normalization condition takes the form:

$$(7.8) \quad \int_0^\infty R^2(r)r^2 dr = 1, \quad \text{or} \quad \int_0^\infty u^2(x) dx = \beta^3, \quad \beta = \frac{mc}{\hbar}.$$

Here, $u = \varphi y$. For further computational details, see our supplementary Mathematica notebook, as well as Refs. [3] and [27]. Final results are presented in Table 5.

In particular, one gets Schrödinger's fine structure formula (for a charged spin-zero particle in the Coulomb field):

$$(7.9) \quad E = E_{n_r} = \frac{mc^2}{\sqrt{1 + \left(\frac{\mu}{n_r + \nu + 1} \right)^2}} \quad (n = n_r = 0, 1, 2, \dots).$$

Here,

$$(7.10) \quad \mu = \frac{Ze^2}{\hbar c}, \quad \nu = -\frac{1}{2} + \sqrt{\left(l + \frac{1}{2} \right)^2 - \mu^2}.$$

The corresponding eigenfunctions are given by the Rodrigues-type formula

$$(7.11) \quad y_n(x) = \frac{B_n}{x^{2\nu+1} e^{-2ax}} \frac{d^n}{dx^n} (x^{n+2\nu+1} e^{-2ax}) = C_n L_n^{2\nu+1}(2ax).$$

Up to a constant, they are Laguerre polynomials (Table 19). In view of the normalization condition (7.8):

$$(7.12) \quad \begin{aligned} \beta^3 &= \int_0^\infty u^2(x) dx = C_n^2 \int_0^\infty [\varphi(x) L_n^{2\nu+1}(2ax)]^2 dx \\ &= \frac{C_n^2}{(2a)^{2\nu+3}} \int_0^\infty e^{-\xi} \xi^{2\nu+2} (L_n^{2\nu+1}(\xi))^2 d\xi, \quad \xi = 2ax. \end{aligned}$$

The corresponding integral is given by (see [41], [42], and Appendix B):

$$(7.13) \quad I_1 = J_{nn1}^{\alpha\alpha} = \int_0^\infty e^{-x} x^{\alpha+1} (L_n^\alpha(x))^2 dx = (\alpha + 2n + 1) \frac{\Gamma(\alpha + n + 1)}{n!}.$$

As a result,

$$(7.14) \quad C_n = 2(a\beta)^{3/2} (2a)^\nu \sqrt{\frac{n!}{\Gamma(2\nu + n + 2)}}.$$

The normalized radial eigenfunctions, corresponding to the relativistic energy levels (7.9), are explicitly given by

$$(7.15) \quad R(r) = R_{n_r}(r) = 2(a\beta)^{3/2} \sqrt{\frac{n_r!}{\Gamma(2\nu + n_r + 2)}} \xi^\nu e^{-\xi/2} L_{n_r}^{2\nu+1}(\xi),$$

where

$$(7.16) \quad \xi = 2ax = 2a\beta r = 2\sqrt{1 - \varepsilon^2} \frac{mc}{\hbar} r.$$

(More details can be found in [3], [7], [39].)

Let us analyze a nonrelativistic limit of Schrödinger's fine structure formula (7.9)–(7.10):

$$(7.17) \quad \begin{aligned} \varepsilon_{\text{Schrödinger}} &= \frac{E_{n_r, l}}{mc^2} = \frac{1}{\sqrt{1 + \frac{\mu^2}{\left(n_r + \frac{1}{2} + \sqrt{\left(l + \frac{1}{2}\right)^2 - \mu^2}\right)^2}}} \\ &= 1 - \frac{\mu^2}{2n^2} - \frac{\mu^4}{2n^4} \left(\frac{n}{l + 1/2} - \frac{3}{4} \right) + O(\mu^6), \quad \mu \rightarrow 0, \end{aligned}$$

which can be derived by a direct Taylor expansion and/or verified by a computer algebra system (see our supplementary Mathematica notebook). Here, $n = n_r + l + 1$ is the corresponding nonrelativistic principal quantum number. The first term in this expansion is simply the rest mass energy $E_0 = mc^2$ of the charged spin-zero particle, the second term coincides with the energy eigenvalue in the nonrelativistic Schrödinger theory and the third term gives the so-called fine structure of the energy levels, which removes the degeneracy between states of the same n and different l .

Once again, our equation,

$$(7.18) \quad xy'' + 2(\nu + 1 - ax)y' + 2(\mu\varepsilon - (\nu + 1)a)y = 0,$$

by the change of variables $y(x) = Y(\xi)$ with $\xi = 2ax$ can be transformed into the required form:

$$(7.19) \quad \xi Y'' + (2\nu + 2 - \xi)Y' \left(\frac{\mu\varepsilon}{\sqrt{1 - \varepsilon^2}} - \nu - 1 \right) Y = 0.$$

TABLE 5. Relativistic Schrödinger equation for Coulomb potential $U(r) = -Ze^2/r$. Dimensionless quantities: $\varepsilon = E/(mc^2)$, $x = \beta r = (mc/\hbar)r$, $\mu = Ze^2/(\hbar c)$, $R(r) = F(x) = u(x)/x$.

$\sigma(x)$	x
$\tilde{\sigma}(x)$	$(\varepsilon^2 - 1)x^2 + 2\mu\varepsilon x + \mu^2 - l(l+1)$
$\tilde{\tau}(x)$	0
k	$2\mu\varepsilon - (2\nu + 1)\sqrt{1 - \varepsilon^2}$, $\nu = -\frac{1}{2} + \sqrt{\left(l + \frac{1}{2}\right)^2 - \mu^2}$
$\pi(x)$	$\nu + 1 - ax$, $a = \sqrt{1 - \varepsilon^2}$
$\tau(x) = \tilde{\tau} + 2\pi$	$2(\nu + 1 - ax)$, $\tau' < 0$
$\lambda = k + \pi'$	$2(\mu\varepsilon - (\nu + 1)a)$
$\varphi(x)$	$x^{\nu+1}e^{-ax}$
$\rho(x)$	$x^{2\nu+1}e^{-2ax}$
$y_{n_r}(x)$	$C_{n_r} L_{n_r}^{2\nu+1}(\xi)$, $\xi = 2ax = 2a\beta r = 2\sqrt{1 - \varepsilon^2} \frac{mc}{\hbar} r$
C_{n_r}	$2(a\beta)^{3/2} (2a)^\nu \sqrt{\frac{n_r!}{(\nu + n_r + 1)\Gamma(2\nu + n_r + 2)}}$

Therefore, the set of the square integrable solutions above is unique.

8. Relativistic Coulomb Problem: Dirac Equation

8.1. System of radial equations. The radial Dirac equations are derived in Refs. [27], [41], and [42] by separation of variables in spherical coordinates (see also [4], [7], [10], [12], and [17]). Then the radial functions $F(r)$ and $G(r)$ satisfy the system of two first-order ordinary differential equations

$$(8.1) \quad \frac{dF}{dr} + \frac{1 + \kappa}{r} F = \frac{mc^2 + E - U(r)}{\hbar c} G,$$

$$(8.2) \quad \frac{dG}{dr} + \frac{1 - \kappa}{r} G = \frac{mc^2 - E + U(r)}{\hbar c} F,$$

where $\kappa = \kappa_\pm = \pm(j + 1/2) = \pm 1, \pm 2, \pm 3, \dots$ respectively. For the relativistic Coulomb problem, when $U = -Ze^2/r$, we introduce the dimensionless quantities

$$(8.3) \quad \varepsilon = \frac{E}{mc^2}, \quad x = \beta r = \frac{mc}{\hbar} r, \quad \mu = \frac{Ze^2}{\hbar c}$$

and change the variable in radial functions

$$(8.4) \quad f(x) = F(r), \quad g(x) = G(r).$$

The Dirac radial system becomes

$$(8.5) \quad \frac{df}{dx} + \frac{1 + \kappa}{x} f = \left(1 + \varepsilon + \frac{\mu}{x}\right) g,$$

$$(8.6) \quad \frac{dg}{dx} + \frac{1 - \kappa}{x} g = \left(1 - \varepsilon - \frac{\mu}{x}\right) f.$$

(One can show later that in the nonrelativistic limit, $c \rightarrow \infty$, the following estimate holds: $|f(x)| \gg |g(x)|$; see, for example, Refs. [27], [41], and [42] for more details.)

8.2. Decoupling of the radial system. We follow [27] with somewhat different details. Let us rewrite the system (8.5)–(8.6) in a matrix form. If

$$(8.7) \quad u = \begin{pmatrix} u_1 \\ u_2 \end{pmatrix} = \begin{pmatrix} xf(x) \\ xg(x) \end{pmatrix}, \quad u' = \begin{pmatrix} u'_1 \\ u'_2 \end{pmatrix}.$$

Then

$$(8.8) \quad u' = Au,$$

where

$$(8.9) \quad A = \begin{pmatrix} a_{11} & a_{12} \\ a_{21} & a_{22} \end{pmatrix} = \begin{pmatrix} -\frac{\kappa}{x} & 1 + \varepsilon + \frac{\mu}{x} \\ 1 - \varepsilon - \frac{\mu}{x} & \frac{\kappa}{x} \end{pmatrix}.$$

To find $u_1(x)$, we eliminate $u_2(x)$ from the system (8.8), obtaining a second-order differential equation

$$(8.10) \quad u_1'' - \left(a_{11} + a_{22} + \frac{a'_{12}}{a_{12}} \right) u_1' + \left(a_{11}a_{22} - a_{12}a_{21} - a'_{11} + \frac{a'_{12}}{a_{12}} a_{11} \right) u_1 = 0.$$

Similarly, eliminating $u_1(x)$, one gets an equation for $u_2(x)$:

$$(8.11) \quad u_2'' - \left(a_{11} + a_{22} + \frac{a'_{21}}{a_{21}} \right) u_2' + \left(a_{11}a_{22} - a_{12}a_{21} - a'_{22} + \frac{a'_{21}}{a_{21}} a_{22} \right) u_2 = 0.$$

The components of the matrix A have the following generic form

$$(8.12) \quad a_{ik} = b_{ik} + c_{ik}/x,$$

where b_{ik} and c_{ik} are constants. Equations (8.10) and (8.11) are not generalized equations of hypergeometric type (2.1). Indeed,

$$\frac{a'_{12}}{a_{12}} = -\frac{c_{12}}{c_{12}x + b_{12}x^2},$$

and the coefficients of $u_1'(x)$ and $u_1(x)$ in (8.10) are

$$a_{11} + a_{22} + \frac{a'_{12}}{a_{12}} = \frac{p_1(x)}{x} - \frac{c_{12}}{c_{12}x + b_{12}x^2},$$

$$a_{11}a_{22} - a_{12}a_{21} - a'_{11} + \frac{a'_{12}}{a_{12}} a_{11} = \frac{p_2(x)}{x^2} - \frac{c_{12}(c_{11} + b_{11}x)}{(c_{12} + b_{12}x)x^2},$$

where $p_1(x)$ and $p_2(x)$ are polynomials of degrees at most one and two, respectively (see supplementary Mathematica notebook for their explicit forms). Equation (8.10) will become a generalized equation of hypergeometric type (2.1) with $\sigma(x) = x$ if either $b_{12} = 0$ or $c_{12} = 0$.

8.3. Similarity transformation. The following consideration helps. By a linear transformation

$$(8.13) \quad \begin{pmatrix} v_1 \\ v_2 \end{pmatrix} = C \begin{pmatrix} u_1 \\ u_2 \end{pmatrix}$$

with a nonsingular matrix C that is independent of x , we transform the original system (8.8) to a similar one

$$(8.14) \quad v' = \tilde{A}v,$$

where

$$v = \begin{pmatrix} v_1 \\ v_2 \end{pmatrix}, \quad \tilde{A} = CAC^{-1} = \begin{pmatrix} \tilde{a}_{11} & \tilde{a}_{12} \\ \tilde{a}_{21} & \tilde{a}_{22} \end{pmatrix}.$$

The new coefficients \tilde{a}_{ik} are linear combinations of the original ones a_{ik} . Hence they have a similar form

$$(8.15) \quad \tilde{a}_{ik} = \tilde{b}_{ik} + \tilde{c}_{ik}/x,$$

where \tilde{b}_{ik} and \tilde{c}_{ik} are constants.

The equations for $v_1(x)$ and $v_2(x)$ are similar to (8.10) and (8.11):

$$(8.16) \quad v_1'' - \left(\tilde{a}_{11} + \tilde{a}_{22} + \frac{\tilde{a}'_{12}}{\tilde{a}_{12}} \right) v_1' + \left(\tilde{a}_{11}\tilde{a}_{22} - \tilde{a}_{12}\tilde{a}_{21} - \tilde{a}'_{11} + \frac{\tilde{a}'_{12}}{\tilde{a}_{12}} \tilde{a}_{11} \right) v_1 = 0,$$

$$(8.17) \quad v_2'' - \left(\tilde{a}_{11} + \tilde{a}_{22} + \frac{\tilde{a}'_{21}}{\tilde{a}_{21}} \right) v_2' + \left(\tilde{a}_{11}\tilde{a}_{22} - \tilde{a}_{12}\tilde{a}_{21} - \tilde{a}'_{22} + \frac{\tilde{a}'_{21}}{\tilde{a}_{21}} \tilde{a}_{22} \right) v_2 = 0.$$

The calculation of the coefficients in (8.16) and (8.17) is facilitated by a similarity of the matrices A and \tilde{A} :

$$\tilde{a}_{11} + \tilde{a}_{22} = a_{11} + a_{22}, \quad \tilde{a}_{11}\tilde{a}_{22} - \tilde{a}_{12}\tilde{a}_{21} = a_{11}a_{22} - a_{12}a_{21}.$$

By a previous consideration, in order for (8.16) to be an equation of hypergeometric type, it is sufficient to choose either $\tilde{b}_{12} = 0$ or $\tilde{c}_{12} = 0$. Similarly, for (8.17): either $\tilde{b}_{21} = 0$ or $\tilde{c}_{21} = 0$. These conditions impose certain restrictions on our choice of the transformation matrix C . Let

$$(8.18) \quad C = \begin{pmatrix} \alpha & \beta \\ \gamma & \delta \end{pmatrix}.$$

Then

$$C^{-1} = \frac{1}{\Delta} \begin{pmatrix} \delta & -\beta \\ -\gamma & \alpha \end{pmatrix}, \quad \Delta = \det C = \alpha\delta - \beta\gamma,$$

and

(8.19)

$$\begin{aligned} \tilde{A} &= CAC^{-1} \\ &= \frac{1}{\Delta} \begin{pmatrix} a_{11}\alpha\delta - a_{12}\alpha\gamma + a_{21}\beta\delta - a_{22}\beta\gamma & a_{12}\alpha^2 - a_{21}\beta^2 + (a_{22} - a_{11})\alpha\beta \\ a_{21}\delta^2 - a_{12}\gamma^2 + (a_{11} - a_{22})\gamma\delta & a_{12}\alpha\gamma - a_{11}\beta\gamma + a_{22}\alpha\delta - a_{21}\beta\delta \end{pmatrix}. \end{aligned}$$

(Here, we have corrected typos in Eqs. (3.74) of [42]; see also [27] and the supplementary Mathematica notebook.) For the Dirac system (8.8)–(8.9):

$$\begin{aligned} a_{11} &= -\frac{\kappa}{x}, & a_{12} &= 1 + \varepsilon + \frac{\mu}{x}, \\ a_{21} &= 1 - \varepsilon - \frac{\mu}{x}, & a_{22} &= \frac{\kappa}{x} \end{aligned}$$

and

$$(8.20) \quad \Delta \tilde{a}_{12} = \alpha^2 - \beta^2 + (\alpha^2 + \beta^2) \varepsilon + \frac{(\alpha^2 + \beta^2) \mu + 2\alpha\beta\kappa}{x},$$

$$(8.21) \quad \Delta \tilde{a}_{21} = \delta^2 - \gamma^2 - (\delta^2 + \gamma^2) \varepsilon - \frac{(\delta^2 + \gamma^2) \mu + 2\gamma\delta\kappa}{x}.$$

The	condition	$\tilde{b}_{12} = 0$	yields	$(1 + \varepsilon)\alpha^2 - (1 - \varepsilon)\beta^2 = 0,$
"	"	$\tilde{c}_{12} = 0$	"	$(\alpha^2 + \beta^2) \mu + 2\alpha\beta\kappa = 0,$
"	"	$\tilde{b}_{21} = 0$	"	$(1 + \varepsilon)\gamma^2 - (1 - \varepsilon)\delta^2 = 0,$
"	"	$\tilde{c}_{21} = 0$	"	$(\delta^2 + \gamma^2) \mu + 2\gamma\delta\kappa = 0.$

We see that there are several possibilities to choose the elements $\alpha, \beta, \gamma, \delta$ of the transition matrix C . All quantum mechanics textbooks use the original one, namely, $\tilde{b}_{12} = 0$ and $\tilde{b}_{21} = 0$, due to Darwin [8] and Gordon [16]. Nikiforov and Uvarov [27] take another path, they choose $\tilde{c}_{12} = 0$ and $\tilde{c}_{21} = 0$ and show that it is more convenient for taking the nonrelativistic limit $c \rightarrow \infty$. These conditions are satisfied if

$$(8.22) \quad C = \begin{pmatrix} \mu & \nu - \kappa \\ \nu - \kappa & \mu \end{pmatrix},$$

where $\nu = \sqrt{\kappa^2 - \mu^2}$, and we finally arrive at the following system of first-order equations for $v_1(x)$ and $v_2(x)$:

$$(8.23) \quad v_1' = \left(\frac{\varepsilon\mu}{\nu} - \frac{\nu}{x} \right) v_1 + \left(1 + \frac{\varepsilon\kappa}{\nu} \right) v_2,$$

$$(8.24) \quad v_2' = \left(1 - \frac{\varepsilon\kappa}{\nu} \right) v_1 + \left(\frac{\nu}{x} - \frac{\varepsilon\mu}{\nu} \right) v_2.$$

Here

$$(8.25) \quad \text{Tr } \tilde{A} = \tilde{a}_{11} + \tilde{a}_{22} = 0, \quad \det \tilde{A} = \varepsilon^2 - 1 + \frac{2\varepsilon\mu}{x} - \frac{\nu^2}{x^2}, \quad \nu^2 = \kappa^2 - \mu^2,$$

which is simpler than the original choice in [27]. The corresponding second-order differential equations (8.16)–(8.17) become

$$(8.26) \quad v_1'' + \frac{(\varepsilon^2 - 1)x^2 + 2\varepsilon\mu x - \nu(\nu + 1)}{x^2} v_1 = 0,$$

$$(8.27) \quad v_2'' + \frac{(\varepsilon^2 - 1)x^2 + 2\varepsilon\mu x - \nu(\nu - 1)}{x^2} v_2 = 0.$$

They are generalized equations of hypergeometric type (2.1) of the simplest form $\tilde{\tau} = 0$, thus resembling the one-dimensional Schrödinger equation; the second equation can be obtained from the first one by replacing $\nu \rightarrow -\nu$ (see also Eqs. (3.81)–(3.82) in Ref. [42]).

8.4. Nikiforov–Uvarov paradigm. All details of the calculations are presented in Table 6 (see also our supplementary Mathematica notebook and Refs. [27], [41], and [42] for more details). Then the corresponding energy levels $\varepsilon = \varepsilon_n$ are determined by

$$(8.28) \quad \varepsilon\mu = a(\nu + n + 1),$$

and the eigenfunctions are given by the Rodrigues-type formula

$$(8.29) \quad y_n(x) = \frac{C_n}{\rho(x)} (\sigma^n(x)\rho(x))^{(n)} = C_n x^{-2\nu-1} e^{2ax} \frac{d^n}{dx^n} (x^{2\nu+n+1} e^{-2ax}).$$

These functions are, up to certain constants, Laguerre polynomials $L_n^{2\nu+1}(\xi)$ (Table 19) with $\xi = 2ax$. The corresponding eigenfunctions have the form

$$(8.30) \quad v_1(x) = \begin{cases} 0, & n = 0, \\ A_n \xi^{\nu+1} e^{-\xi/2} L_{n-1}^{2\nu+1}(\xi), & n = 1, 2, 3, \dots \end{cases}$$

They are square integrable functions on $(0, \infty)$. The counterparts are

$$(8.31) \quad v_2(x) = B_n \xi^\nu e^{-\xi/2} L_n^{2\nu-1}(\xi), \quad n = 0, 1, 2, \dots$$

It is easily seen that the solution $\varepsilon = -\nu/\kappa$ is included in this formula when $n = 0$.

As a result,

$$(8.32) \quad xf(x) = \frac{B_n}{2\nu(\kappa - \nu)} \xi^\nu e^{-\xi/2} (f_1 \xi L_{n-1}^{2\nu+1}(\xi) + f_2 L_n^{2\nu-1}(\xi)),$$

$$(8.33) \quad xg(x) = \frac{B_n}{2\nu(\kappa - \nu)} \xi^\nu e^{-\xi/2} (g_1 \xi L_{n-1}^{2\nu+1}(\xi) + g_2 L_n^{2\nu-1}(\xi)),$$

where

$$(8.34) \quad f_1 = \frac{a\mu}{\varepsilon\kappa - \nu}, \quad f_2 = \kappa - \nu, \quad g_1 = \frac{a(\kappa - \nu)}{\varepsilon\kappa - \nu}, \quad g_2 = \mu.$$

(These formulas remain valid for $n = 0$; in this case the terms containing $L_{-1}^{2\nu+1}(\xi)$ have to be taken to be zero.) Thus we derive the representation for the radial functions up to the constant B_n in terms of Laguerre polynomials (Table 19). The normalization condition

$$(8.35) \quad \int_{\mathbb{R}^3} \psi^\dagger \psi \, dv = \int_0^\infty r^2 (F^2(r) + G^2(r)) \, dr = 1$$

gives the value of this constant as follows [27]:

$$(8.36) \quad B_n = a\beta^{3/2} \sqrt{\frac{(\kappa - \nu)(\varepsilon\kappa - \nu)n!}{\mu\Gamma(n + 2\nu)}}.$$

(This is verified in section 5.4 of Ref. [42]. Observe that Eq. (8.36) applies when $n = 0$.)

8.5. Summary: wave functions and energy levels. The end results, namely, the complete wave functions and the corresponding discrete energy levels, are given by Eqs. (3.11)–(3.17) of Ref. [42]. The WKB, or semiclassical, approximation for the Dirac equation with Coulomb potential is discussed in [3].

The relativistic energy levels of an electron in the central Coulomb field are given by

$$(8.37) \quad E = E_{n_r, j} = \frac{mc^2}{\sqrt{1 + \mu^2/(n_r + \nu)^2}}, \quad \mu = \frac{Ze^2}{\hbar c} \quad (n_r = 0, 1, 2, \dots).$$

In Dirac's theory,

$$(8.38) \quad \nu = \nu_{\text{Dirac}} = \sqrt{(j + 1/2)^2 - \mu^2},$$

where $j = 1/2, 3/2, 5/2, \dots$ is the total angular momentum including the spin of the relativistic electron. More details on the solution of this problem, including the nonrelativistic limit, can be found in [27], [41], [42] (following Nikiforov–Uvarov's paradigm), or in classical sources [4], [7], [8], [12], [16], [39].

In Dirac's theory of the relativistic electron, the corresponding limit has the form [4], [7], [39], [42]:

$$(8.39) \quad \varepsilon_{\text{Dirac}} = \frac{E_{n_r, j}}{mc^2} = 1 - \frac{\mu^2}{2n^2} - \frac{\mu^4}{2n^4} \left(\frac{n}{j + 1/2} - \frac{3}{4} \right) + O(\mu^6), \quad \mu \rightarrow 0,$$

where $n = n_r + j + 1/2$ is the principal quantum number of the nonrelativistic hydrogenlike atom. Once again, the first term in this expansion is the rest mass energy of the relativistic electron, the second term coincides with the energy eigenvalue in the nonrelativistic Schrödinger theory and the third term gives the so-called fine structure of the energy levels — the correction obtained for the energy in the Pauli approximation which includes the interaction of the spin of the electron with its orbital angular momentum. (See our supplementary Mathematica notebook for a computer algebra proof.)

9. A Model of the 3D-confinement Potential

Looking for solutions of the Schrödinger equation (5.1) in spherical coordinates,

$$(9.1) \quad \psi = \frac{1}{r} R(r) Y_{lm}(\theta, \varphi),$$

with the following model central field potential,

$$(9.2) \quad U(r) = V_0 \left(\frac{r}{a} - \frac{a}{r} \right)^2,$$

one gets the radial equation of the form

$$(9.3) \quad R'' + \frac{2m}{\hbar^2} \left[(E + 2V_0) - V_0 \left(\frac{a^2}{r^2} + \frac{r^2}{a^2} \right) - \frac{\hbar^2 l(l+1)}{2mr^2} \right] R = 0 \quad (l = 0, 1, 2, \dots).$$

TABLE 6. Dirac equation for the Coulomb potential $U(r) = -Ze^2/r$. Dimensionless quantities: $\varepsilon = E/(mc^2)$, $x = \beta r = (mc/\hbar)r$, $\mu = Ze^2/(\hbar c)$, $F(r) = f(x) = u_1(x)/x$, $G(r) = g(x) = u_2(x)/x$.

$\sigma(x)$	x
$\tilde{\sigma}(x)$	$(\varepsilon^2 - 1)x^2 + 2\mu\varepsilon x - \nu(\nu + 1); \nu = \sqrt{\kappa^2 - \mu^2}, \kappa = \pm \left(j + \frac{1}{2}\right)$
$\tilde{\tau}(x)$	0
k	$2\mu\varepsilon \pm \sqrt{1 - \varepsilon^2}(2\nu + 1)$
$\pi(x)$	$\frac{1}{2} \pm \left(\sqrt{1 - \varepsilon^2} x \pm \left(\nu + \frac{1}{2}\right)\right)$
$\tau(x) = \tilde{\tau} + 2\pi$	$2(\nu + 1 - ax), \quad a = \sqrt{1 - \varepsilon^2}; \quad \tau' < 0$
$\lambda = k + \pi'$	$2(\mu\varepsilon - (\nu + 1)a)$
$\varphi(x)$	$x^{\nu+1}e^{-ax}$
$\rho(x)$	$x^{2\nu+1}e^{-2ax}$
$y_n(x)$	$A_n L_n^{2\nu+1}(\xi); \quad \xi = 2ax = 2a\beta r = 2\sqrt{1 - \varepsilon^2} \frac{mc}{\hbar} r$ $n = n_r = 0, 1, \dots$
$A_n = \frac{a}{\kappa\varepsilon - \nu} B_n$	$B_n = a\beta^{3/2} \sqrt{\frac{(\kappa - \nu)(\kappa\varepsilon - \nu)n!}{\mu\Gamma(2\nu + n)}}$

This is not a generalized equation of hypergeometric type and, therefore, cannot be treated right away by the Nikiforov–Uvarov method. By using the substitution

$$(9.4) \quad R(r) = u(\xi), \quad \xi = \alpha r^2, \quad \alpha^2 = \frac{2mV_0}{\hbar^2 a^2},$$

we finally obtain equation (2.1) with the following coefficients:

$$(9.5) \quad \sigma(\xi) = \xi, \quad \tilde{\tau} = \frac{1}{2}, \quad \tilde{\sigma}(\xi) = \frac{1}{4} \left[\frac{2m}{\alpha\hbar^2} (E + 2V_0)\xi - \alpha^2 a^4 - l(l+1) - \xi^2 \right].$$

In the Nikiforov–Uvarov method, the energy levels and the corresponding radial wave functions can be obtained by (2.7) and (2.8). As a result, they are given by

$$(9.6) \quad E_{n,l} = \hbar \sqrt{\frac{8V_0}{ma^2}} \left[n + \frac{1}{2} + \frac{1}{2} \left(\sqrt{\frac{2mV_0 a^2}{\hbar^2} + \left(l + \frac{1}{2}\right)^2} - \sqrt{\frac{2mV_0 a^2}{\hbar^2}} \right) \right]$$

and

$$(9.7) \quad R(r) = R_{n,l}(r) = C_n \xi^{(\beta-1/2)/2} \exp(-\xi/2) L_n^\beta(\xi),$$

provided

$$(9.8) \quad \int_0^\infty R^2(r) dr = 1,$$

respectively. Here

$$(9.9) \quad C_n^2 = \frac{2n!\sqrt{\alpha}}{\Gamma(\beta + n + 1)}, \quad \alpha = \frac{\sqrt{2mV_0}}{\hbar a}, \quad \beta = \sqrt{\frac{2mV_0a^2}{\hbar^2} + \left(l + \frac{1}{2}\right)^2}.$$

Details of the calculations are presented in Table 7. (The case $l = 0$ corresponds to a one-dimensional problem from [14].)

In this case, the spectrum is linear, as for the harmonic oscillator. There is no continuous spectrum thus resembling the confinement property in quantum chromodynamics.

Energy levels (9.6) can be expanded as follows

$$(9.10) \quad E_{n,l} = \hbar\omega \left(n + \frac{1}{2}\right) + \frac{\hbar^2}{2ma^2} \left(l + \frac{1}{2}\right)^2 - \frac{\hbar^4}{4aV_0(2m)^{3/2}} \left(l + \frac{1}{2}\right)^4 + \dots, \quad \frac{(l + 1/2)^2}{2maV_0} \ll 1.$$

Here

$$(9.11) \quad \omega = \sqrt{\frac{8V_0}{ma^2}},$$

the first term represents the familiar harmonic oscillations, the second one is the rotational energy, and the last term gives its correction in the nonsymmetric well.

10. 3D-Spherical Oscillator

Looking for solutions of the Schrödinger equation (5.1) with harmonic potential

$$(10.1) \quad U(r) = \frac{1}{2}m\omega^2r^2$$

in spherical coordinates (9.1), one gets the following radial equation:

$$(10.2) \quad R'' + \left[\frac{2mE}{\hbar^2} - \frac{m^2\omega^2}{\hbar^2}r^2 - \frac{l(l+1)}{r^2} \right] R = 0 \quad (l = 0, 1, 2, \dots).$$

Using the abbreviations [10]

$$(10.3) \quad \frac{2mE}{\hbar^2} = \kappa^2, \quad \frac{m\omega}{\hbar} = \mu, \quad \frac{\kappa^2}{2\mu} = \frac{E}{\hbar\omega} = \varepsilon,$$

the radial equation can be rewritten in the standard form

$$(10.4) \quad \frac{d^2R}{dr^2} + \left[\kappa^2 - \mu^2r^2 - \frac{l(l+1)}{r^2} \right] R = 0.$$

Finally, the substitution $R(r) = u(\xi)$ with $\xi = \mu r^2$ results in the generalized equation of hypergeometric type with

$$(10.5) \quad \sigma(\xi) = \xi, \quad \tilde{\tau}(\xi) = \frac{1}{2}, \quad \tilde{\sigma}(\xi) = \frac{1}{4} \left[\frac{\kappa^2}{\mu} \xi - \xi^2 - l(l+1) \right].$$

Therefore,

$$(10.6) \quad 2k - \varepsilon = \pm \left(l + \frac{1}{2} \right), \quad \pi(\xi) = \frac{1}{4} \pm \frac{1}{2} \left[\xi \pm \left(l + \frac{1}{2} \right) \right].$$

Further details of calculation are presented in Table 8 and in the corresponding Mathematica file.

TABLE 7. A model of the 3D-confinement potential $U(r)$ given in (9.2). Dimensionless quantities: $\xi = \alpha r^2$, $\alpha = \sqrt{2mV_0}/(\hbar a)$, $R(r) = u(\xi)$.

$\sigma(\xi)$	ξ
$\tilde{\sigma}(\xi)$	$\frac{1}{4} \left[\frac{2m}{\alpha \hbar^2} (E + 2V_0) \xi - \alpha^2 a^4 - l(l+1) - \xi^2 \right]$
$\tilde{\tau}(\xi)$	$1/2$
k	$\frac{1}{2} \left[\frac{m}{\alpha \hbar^2} (E + 2V_0) \pm \sqrt{\alpha^2 a^4 + (l+1/2)^2} \right]$
$\pi(\xi)$	$\frac{1}{4} \pm \frac{\xi \pm \sqrt{\alpha^2 a^4 + (l+1/2)^2}}{2}$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$1 + \beta - \xi, \quad \beta = \sqrt{\frac{2mV_0 a^2}{\hbar^2} + \left(l + \frac{1}{2}\right)^2}$
$\lambda = k + \pi'$	$\frac{1}{2} \left[\frac{m}{\alpha \hbar^2} (E + 2V_0) - \sqrt{\alpha^2 a^4 + (l+1/2)^2} \right] - \frac{1}{2}$
$\varphi(\xi)$	$\xi^{\nu/2} \exp(-\xi/2), \quad \nu = \beta + \frac{1}{2} = \frac{1}{2} + \sqrt{\alpha^2 a^4 + \left(l + \frac{1}{2}\right)^2}$
$\rho(\xi)$	$\xi^\beta \exp(-\xi)$
$y_n(\xi)$	$C_n L_n^\beta(\xi), \quad \xi = \alpha r^2, \quad n = 0, 1, \dots$
C_n^2	$\frac{2n! \sqrt{\alpha}}{\Gamma(\beta + n + 1)}, \quad \alpha = \frac{\sqrt{2mV_0}}{\hbar a}$

As a result, the energy levels are given by

$$(10.7) \quad E_n = \hbar\omega \left(2n + l + \frac{3}{2} \right), \quad n = 0, 1, 2, \dots$$

and the corresponding radial wave functions are related to the Laguerre polynomials (Table 19):

$$(10.8) \quad R_n(r) = C_n \xi^{l+1} \exp\left(-\frac{\xi}{2}\right) L_n^{l+1/2}(\xi), \quad n = 0, 1, 2, \dots$$

Here,

$$(10.9) \quad \int_0^\infty R_n^2(r) dr = 1.$$

Extension to the case of n -dimensions is discussed in [26].

11. Pöschl–Teller Potential Hole

Let us consider the one-dimensional stationary Schrödinger equation:

$$(11.1) \quad -\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} + U(x)\psi = E\psi,$$

TABLE 8. Data for the 3D-spherical harmonic oscillator potential $U(x) = \frac{1}{2}m\omega^2 r^2$. Substitution: $\xi = \mu r^2$, $\mu = m\omega/\hbar$, and $R(r) = u(\xi)$.

$\sigma(\xi)$	ξ
$\tilde{\sigma}(\xi)$	$\frac{1}{4} [2\varepsilon \xi - \xi^2 - l(l+1)]$, $\varepsilon = \frac{E}{\hbar\omega}$
$\tilde{\tau}(\xi)$	$\frac{1}{2}$
k	$-\frac{1}{2} \left(l + \frac{1}{2} - \varepsilon \right)$
$\pi(\xi)$	$\frac{1}{2}(l+1-\xi)$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$l + \frac{3}{2} - \xi$
$\lambda = k + \pi'$	$-\frac{1}{2} \left(l + \frac{3}{2} - \varepsilon \right)$
$\varphi(\xi)$	$\xi^{(l+1)/2} \exp(-\xi/2)$
$\rho(\xi)$	$\xi^{l+1/2} \exp(-\xi)$
$y_n(\xi)$	$C_n L_n^{l+1/2}(\xi)$, $n = 0, 1, \dots$
C_n^2	$\frac{2n! \sqrt{\mu}}{\Gamma(l+n+3/2)}$, $\mu = \frac{m\omega}{\hbar}$

where

$$(11.2) \quad U(x) = \frac{1}{2}V_0 \left[\frac{a(a-1)}{\sin^2(\alpha x)} + \frac{b(b-1)}{\cos^2(\alpha x)} \right], \quad V_0 = \frac{\hbar^2 \alpha^2}{m}$$

with real-valued parameters $a > 1$, $b > 1$ in the finite region $0 < x < \pi/(2\alpha)$ bounded by the singularities of $U(x)$ (see [10], [30], [32] for original references and applications). Here, we are looking for orthonormal real-valued wave functions:

$$(11.3) \quad \int_0^{\pi/(2\alpha)} \psi^2(x) dx = 1.$$

Introducing new quantities

$$(11.4) \quad \psi(x) = u(\xi), \quad \xi = \sin^2(\alpha x), \quad 1 - \xi = \cos^2(\alpha x),$$

one gets the following generalized equation of hypergeometric type:

$$(11.5) \quad \xi(1-\xi)u'' + \left(\frac{1}{2} - \xi \right) u' + \frac{1}{4} \left[\frac{c^2}{\alpha^2} - \frac{a(a-1)}{\xi} - \frac{b(b-1)}{1-\xi} \right] u = 0, \quad c^2 = \frac{2mE}{\hbar^2}.$$

Here,

$$(11.6) \quad \sigma(\xi) = \xi(1 - \xi), \quad \tilde{\tau}(\xi) = \left(\frac{1}{2} - \xi\right),$$

$$\tilde{\sigma}(\xi) = \frac{1}{4} \left[\frac{c^2}{\alpha^2} \xi(1 - \xi) - a(a - 1)(1 - \xi) - b(b - 1)\xi \right]$$

and the boundary conditions take the form $u(0) = u(1) = 0$.

Therefore,

$$(11.7) \quad p(\xi) = \frac{1}{4} \left[(\kappa + 1)\xi^2 - (\kappa + 1 + (a - b)(a + b - 1))\xi + a(a - 1) + \frac{1}{4} \right],$$

$$\kappa = \frac{c^2}{\alpha^2} - 4k.$$

Equation (2.21) takes the form

$$(11.8) \quad [\kappa + 1 + (a - b)(a + b - 1)]^2 = (\kappa + 1) [4a(a - 1) + 1].$$

There are two solutions

$$(11.9) \quad \kappa_1 = (a + b)(a + b - 2), \quad \kappa_2 = -(b - a - 1)(b - a + 1)$$

If one chooses

$$(11.10) \quad \frac{c^2}{\alpha^2} - 4k = (a + b)(a + b - 2),$$

then

$$(11.11) \quad \pi(\xi) = \frac{1}{2} \left(\frac{1}{2} - \xi \right) - \left(\frac{a + b - 1}{2} \xi - \frac{(2a - 1)}{4} \right) = \frac{a}{2} - \frac{a + b}{2} \xi$$

and

$$(11.12) \quad \tau(\xi) = a + \frac{1}{2} - (a + b + 1)\xi, \quad \lambda = \frac{1}{4} \left[\frac{c^2}{\alpha^2} - (a + b)^2 \right].$$

Further details of calculation are presented in Table 9 (see also the corresponding Mathematica file).

As a result, the energy levels are given by (2.7):

$$(11.13) \quad E_n = \frac{1}{2} V_0 (a + b + 2n)^2, \quad n = 0, 1, 2, \dots$$

and the corresponding wave functions are related to the Jacobi polynomials:

$$(11.14) \quad \psi_n(x) = C_n \sin^a(\alpha x) \cos^b(\alpha x) P_n^{(a-1/2, b-1/2)}(\cos(2\alpha x)),$$

where C_n is the normalization constant.

Indeed, by the Rodrigues-type formula (2.8):

$$(11.15) \quad y_n = \frac{B_n}{\xi^{a-1/2}(1 - \xi)^{b-1/2}} \frac{d^n}{d\xi^n} \left[\xi^{n+a-1/2}(1 - \xi)^{n+b-1/2} \right]$$

and with the aid of the substitution $\eta = 1 - 2\xi = \cos(2\alpha x)$ one gets

$$(11.16) \quad y_n(\eta) = \frac{(-1)^n}{2^n} \frac{B_n}{(1 - \eta)^{a-1/2}(1 + \eta)^{b-1/2}} \frac{d^n}{d\eta^n} \left[(1 - \eta)^{n+a-1/2}(1 + \eta)^{n+b-1/2} \right]$$

$$= C_n P_n^{(a-1/2, b-1/2)}(\eta).$$

TABLE 9. The Pöschl–Teller potential hole $U(x)$ is given in (11.2).
 Substitution: $\xi = \sin^2(\alpha x)$, $\psi(x) = u(\xi)$, $0 < x < \pi/(2\alpha)$.

$\sigma(\xi)$	$\xi(1 - \xi) = \xi - \xi^2$
$\tilde{\sigma}(\xi)$	$-\frac{1}{4} \left[\frac{c^2}{\alpha^2} \xi^2 - \left(\frac{c^2}{\alpha^2} + (a - b)(a + b - 1) \right) \xi + a(a - 1) \right]$, $c^2 = \frac{2mE}{\hbar^2}$
$\tilde{\tau}(\xi)$	$\frac{1}{2} - \xi$
k	$\frac{1}{4} \left[\frac{c^2}{\alpha^2} - (a + b)(a + b - 2) \right]$
$\pi(\xi)$	$\frac{1}{2} [a - (a + b)\xi]$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$(a + 1/2) - (a + b + 1)\xi$
$\lambda = k + \pi'$	$\frac{1}{4} \left[\frac{c^2}{\alpha^2} - (a + b)^2 \right]$
$\varphi(\xi)$	$\xi^{a/2}(1 - \xi)^{b/2} = \sin^a(\alpha x) \cos^b(\alpha x)$
$\rho(\xi)$	$\xi^{a-1/2}(1 - \xi)^{b-1/2}$
$y_n(\xi)$	$C_n P_n^{(a-1/2, b-1/2)}((1 - \xi)/2) = C_n P_n^{(a-1/2, b-1/2)}(\cos(2\alpha x))$
C_n^2	$\frac{2\alpha n!(a + b + 2n)\Gamma(a + b + n)}{\Gamma(a + n + 1/2)\Gamma(b + n + 1/2)} \quad (n = 0, 1, 2, \dots)$

Moreover, by the normalization condition:

$$\begin{aligned}
 (11.17) \quad 1 &= \int_0^{\pi/(2\alpha)} \psi_n^2(x) dx = \frac{C_n^2}{2\alpha} \int_{-1}^1 \left[P_n^{(a-1/2, b-1/2)}(\eta) \right]^2 \frac{\varphi^2 d\eta}{(1 - \eta)^{1/2}(1 + \eta)^{1/2}} \\
 &= \frac{C_n^2}{(2\alpha)2^{a+b}} \int_{-1}^1 \left[P_n^{(a-1/2, b-1/2)}(\eta) \right]^2 (1 - \eta)^{a-1/2}(1 + \eta)^{b-1/2} d\eta \\
 &= \frac{C_n^2}{2\alpha} \frac{\Gamma(a + n + 1/2)\Gamma(b + n + 1/2)}{n!(a + b + 2n)\Gamma(a + b + n)}.
 \end{aligned}$$

(Here, the value of the squared norm d_n^2 for the Jacobi polynomials has been taken from Table 19.) As a result,

$$(11.18) \quad C_n^2 = \frac{2\alpha n!(a + b + 2n)\Gamma(a + b + n)}{\Gamma(a + n + 1/2)\Gamma(b + n + 1/2)}.$$

12. Modified Pöschl–Teller Potential Hole

In order to solve the one-dimensional stationary Schrödinger equation (11.1) for the potential:

$$(12.1) \quad U(x) = -\frac{\hbar^2 \alpha^2}{2m} \frac{a(a - 1)}{\cosh^2(\alpha x)} \quad (-\infty < x < \infty)$$

with $a > 1$, one can use the following substitution $\psi(x) = u(\xi)$, where $\xi = \cosh^2(\alpha x)$ [10], [30]. As a result, we arrive at the generalized equation of hypergeometric type

$$(12.2) \quad \xi(1-\xi)u'' + \left(\frac{1}{2} - \xi\right)u' + \frac{1}{4} \left[\frac{c^2}{\alpha^2} - \frac{a(a-1)}{\xi} \right] u = 0, \quad c^2 = \frac{2m(-E)}{\hbar^2},$$

where

$$(12.3) \quad \begin{aligned} \sigma(\xi) &= \xi(1-\xi), & \tilde{\tau}(\xi) &= \left(\frac{1}{2} - \xi\right), \\ \tilde{\sigma}(\xi) &= \frac{1}{4} \left[\frac{c^2}{\alpha^2} \xi - a(a-1) \right] (1-\xi). \end{aligned}$$

Using the standard substitution $u = \varphi(\xi)y$ with $\varphi(\xi) = \xi^{(1-a)/2}$, one gets the hypergeometric differential equation of the form:

$$(12.4) \quad \xi(1-\xi)y'' + \left[\frac{3}{2} - a + (a-2)\xi \right] y' + \frac{1}{4} \left[\frac{c^2}{\alpha^2} - (a-1)^2 \right] y = 0.$$

Here, we concentrate only on the bounded states (continuous spectrum is discussed in [10]).

There is a finite number of negative discrete energy levels that are explicitly given by

$$(12.5) \quad E = E_n = -\frac{\hbar^2 \alpha^2}{2m} (1-a+2n)^2, \quad n = 0, 1, \dots < (a-1)/2.$$

The corresponding orthonormal wave functions are related to a set of Jacobi polynomials with a negative value of one parameter that are orthogonal on an infinite interval $(1, +\infty)$. They are given by the Rodrigues-type formula (2.8) or in terms of a terminating hypergeometric series:

$$(12.6) \quad \begin{aligned} y_n(\xi) &= \frac{(1/2)_n}{n!} {}_2F_1 \left(-n, 1-a+n; \frac{1}{2}; 1-\xi \right) \\ &= P_n^{(-1/2, 1/2-a)}(\cosh(2\alpha x)), \end{aligned}$$

where $(1/2)_n = \Gamma(n+1/2)/\Gamma(1/2)$. Cauchy's beta integral,

$$(12.7) \quad \int_0^\infty \frac{t^{A-1}}{(1+t)^{A+B}} dt = \frac{\Gamma(A)\Gamma(B)}{\Gamma(A+B)}, \quad \Re(A) > 0, \Re(B) > 0,$$

should be used in order to find the value of the normalization constant (see [42], Exercise 1.15 and [28], (5.12.3)). Further details of calculation are presented in Table 10 (see also the corresponding Mathematica file). As a result, for the bound states (12.5), the normalized wave functions are given by

$$(12.8) \quad \begin{aligned} \psi_n(x) &= \sqrt{\alpha} \left[\frac{n!(a-1-2n)\Gamma(a-n-1/2)}{\Gamma(n+1/2)\Gamma(a-n)} \right]^{1/2} \\ &\quad \times \cosh^{1-a}(\alpha x) P_n^{(-1/2, 1/2-a)}(\cosh(2\alpha x)). \end{aligned}$$

TABLE 10. The modified Pöschl–Teller potential hole $U(x)$ is defined in (12.1). Substitution: $\xi = \cosh^2(\alpha x)$, $\psi(x) = u(\xi)$.

$\sigma(\xi)$	$\xi(1 - \xi)$
$\tilde{\sigma}(\xi)$	$\frac{1}{4} \left[\frac{c^2}{\alpha^2} \xi - a(a - 1) \right] (1 - \xi), \quad c^2 = \frac{2m(-E)}{\hbar^2}$
$\tilde{\tau}(\xi)$	$\frac{1}{2} - \xi$
k	$\frac{1}{4} \left(\frac{c^2}{\alpha^2} + 1 - a^2 \right)$
$\pi(\xi)$	$\frac{1 - a}{2} (1 - \xi)$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$\frac{3}{2} - a + (a - 2)\xi$
$\lambda = k + \pi'$	$\frac{1}{4} \left[\frac{c^2}{\alpha^2} - (a - 1)^2 \right]$
$\varphi(\xi)$	$\xi^{(1-a)/2}$
$\rho(\xi)$	$\xi^{(1/2)-a} (\xi - 1)^{-(1/2)}$
$y_n(\xi)$	$C_n P_n^{(-1/2, 1/2-a)}(\cosh(2\alpha x)), \quad n < (a - 1)/2$
C_n^2	$\alpha \frac{n!(a - 1 - 2n)\Gamma(a - n - 1/2)}{\Gamma(n + 1/2)\Gamma(a - n)}$

13. Kratzer's Molecular Potential

In order to investigate the rotation-vibration spectrum of a diatomic molecule, the potential

$$(13.1) \quad U(r) = -2D \left(\frac{a}{r} - \frac{1}{2} \frac{a^2}{r^2} \right) \quad D > 0,$$

with a minimum $U(a) = -D$, has been used [10]. Once again we are looking for solutions of the Schrödinger equation (5.1) in spherical coordinates (9.1) and introduce the dimensionless quantities:

$$(13.2) \quad x = \frac{r}{a}, \quad \beta^2 = -\frac{2ma^2}{\hbar^2} E, \quad \gamma^2 = \frac{2ma^2}{\hbar^2} D$$

together with the standard substitution: $R(r) = u(x)$.

For bound states $E < 0, \beta > 0$ and the radial equation takes the form

$$(13.3) \quad u'' + \left[-\beta^2 + \frac{2\gamma^2}{x} - \frac{\gamma^2 + l(l + 1)}{x^2} \right] u = 0.$$

Further computational details are presented in Table 11 (see also the corresponding Mathematica file). This case is somewhat similar to Coulomb and relativistic Coulomb problems.

As a result, the bound states are given by

$$(13.4) \quad E_{n,l} = -\frac{2ma^2D^2}{\hbar^2} \frac{1}{(\nu+n)^2},$$

where

$$(13.5) \quad \nu = \frac{1}{2} + \sqrt{\gamma^2 + \left(l + \frac{1}{2}\right)^2}, \quad \gamma^2 = \frac{2ma^2}{\hbar^2} D.$$

We can obtain the same exact result with the aid of the Bohr–Sommerfeld quantization rule in the semiclassical approximation (the WKB-method [3], [22], [27], [40]).

The dimensionless parameter γ^2 is proportional to the reduced mass of the nuclei m . When $\gamma \gg 1$ and $n \ll \gamma$, $l \ll \gamma$, the energy (13.4) can be expanded as follows

$$(13.6) \quad \begin{aligned} E_{n,l} = & -D + \hbar\omega \left(n + \frac{1}{2}\right) + \frac{\hbar^2}{2ma^2} \left(l + \frac{1}{2}\right)^2 \\ & - \frac{2\hbar^2}{ma^2} \left[\frac{3}{4} \left(n + \frac{1}{2}\right)^2 - \frac{1}{\gamma} \left(n + \frac{1}{2}\right)^3 \right] \\ & - \frac{3\hbar^2}{2m\gamma a^2} \left(n + \frac{1}{2}\right) \left(l + \frac{1}{2}\right)^2 + O(\gamma^{-4}), \quad \gamma \gg 1 \end{aligned}$$

(see our Mathematica notebook). Here

$$(13.7) \quad \hbar\omega = \frac{2D}{\gamma} = \frac{\hbar}{a} \sqrt{\frac{2D}{m}}.$$

The second and third terms in this expression give the energy of the vibrational and rotational motion, respectively. The fourth term takes into account anharmonic oscillations, and, finally, the last term gives the correction to the energy due to the interaction between the rotational and vibrational motions of the nuclei [10], [14] (we have corrected a couple of typos therein).

The normalized radial wave functions,

$$(13.8) \quad \int_0^\infty R^2(r) dr = 1,$$

are related to the Laguerre polynomials (Table 19):

$$(13.9) \quad R_n(r) = C_n x^\nu \exp(-\beta x) L_n^{2\nu-1}(2\beta x) \quad \left(x = \frac{r}{a}\right),$$

where

$$(13.10) \quad C_n^2 = \frac{(2\beta)^{2\nu+1} n!}{a(2\nu+2n)\Gamma(2\nu+n)}.$$

Once again, we have used the integral (7.13). (See also [1], [5], [10], [13], [14], [18] and the Mathematica file for some applications of Kratzer’s molecular potential, graphs, and further references.)

TABLE 11. Kratzer's molecular potential (13.1). Substitution:
 $x = r/a$, $R(r) = u(x)$; see also (13.2).

$\sigma(x)$	x
$\tilde{\sigma}(x)$	$2\gamma^2 x - \beta^2 x^2 - \gamma^2 - l(l+1)$
$\tilde{\tau}(x)$	0
k	$2\gamma^2 \pm \beta(2\nu - 1), \quad \nu = \frac{1}{2} + \sqrt{\gamma^2 + \left(l + \frac{1}{2}\right)^2}$
$\pi(x)$	$\nu - \beta x$
$\tau(x) = \tilde{\tau} + 2\pi$	$2(\nu - \beta x)$
$\lambda = k + \pi'$	$2(\gamma^2 - \nu\beta)$
$\varphi(x)$	$x^\nu \exp(-\beta x)$
$\rho(x)$	$x^{2\nu-1} \exp(-2\beta x)$
$y_n(x)$	$C_n L_n^{2\nu-1}(2\beta x)$
C_n^2	$\frac{(2\beta)^{2\nu+1} n!}{a(2\nu + 2n)\Gamma(2\nu + n)}, \quad \beta = \beta_n = \frac{2ma^2}{\hbar^2} \frac{D}{\nu + n}$

14. Hulthén Potential

We are looking for solutions of the Schrödinger equation (5.1) in spherical coordinates (9.1) for the following central field potential,

$$(14.1) \quad U(r) = -V_0 \frac{e^{-r/a}}{1 - e^{-r/a}} \quad (0 \leq r < \infty),$$

when $l = 0$. (See, for example, (6.2) with $F = R$ and use these data for an explicit form of the corresponding radial equation.)

With the aid of the substitution

$$(14.2) \quad R(r) = u(\xi), \quad \xi = e^{-r/a},$$

when

$$(14.3) \quad \alpha^2 = -\frac{2mE}{\hbar^2} a^2 > 0, \quad \beta^2 = \frac{2mV_0}{\hbar^2} a^2 > 0$$

with $\alpha > 0$ and $\beta > 0$, one obtains the following generalized equation of hypergeometric type,

$$(14.4) \quad \xi^2 \frac{d^2 u}{d\xi^2} + \xi \frac{du}{d\xi} + \left(-\alpha^2 + \beta^2 \frac{\xi}{1-\xi} \right) u = 0,$$

where

$$(14.5) \quad \sigma(\xi) = \xi(1-\xi), \quad \tilde{\tau}(\xi) = 1-\xi,$$

and

$$(14.6) \quad \tilde{\sigma}(\xi) = (1-\xi) \left((\alpha^2 + \beta^2) \xi - \alpha^2 \right).$$

The boundary conditions are

$$(14.7) \quad R(0) = \lim_{r \rightarrow 0^+} R(r) = 0, \quad R(\infty) = \lim_{r \rightarrow \infty} R(r) = 0,$$

or

$$(14.8) \quad u(0) = u(1) = 0.$$

We have

$$(14.9) \quad k = \beta^2 \mp \alpha, \quad \pi(\xi) = -\frac{\xi}{2} \pm \left(-\alpha \pm \frac{\xi}{2} + \alpha\xi \right)$$

for all four possible sign combinations. The solutions can be found by putting [10]

$$(14.10) \quad u = \varphi(\xi)y(\xi) = \xi^\alpha(1-\xi)y(\xi),$$

which results in

$$(14.11) \quad \xi(1-\xi)y'' + [2\alpha + 1 - (2\alpha + 3)\xi]y' - (2\alpha + 1 - \beta^2)y = 0$$

(details of calculations are presented in Table 12 and in a complementary Mathematica file).

As one can see, a direct quantization in terms of the classical orthogonal polynomials by Nikiforov and Uvarov's approach is not applicable here, right away, because the second coefficient,

$$(14.12) \quad \tau(\xi) = 2\alpha + 1 - (2\alpha + 3)\xi,$$

does depend on α and therefore on the energy E . We have to utilize the boundary conditions (14.8) instead, in a somewhat similar way to the consideration of a familiar case of an infinite well.

Equation (14.11) is a special case of the hypergeometric equation [2], [28],

$$(14.13) \quad \xi(1-\xi)w'' + [C - (A + B + 1)\xi]w' - ABw = 0,$$

with

$$(14.14) \quad A = 1 + \alpha + \gamma, \quad B = 1 + \alpha - \gamma, \quad C = 2\alpha + 1; \quad \gamma = \sqrt{\alpha^2 + \beta^2}.$$

The required solution, that is bounded at $\xi = 0$, has the form

$$(14.15) \quad y = {}_2F_1\left(\begin{matrix} A, B \\ C \end{matrix}; \xi\right) = {}_2F_1\left(\begin{matrix} 1 + \alpha + \gamma, 1 + \alpha - \gamma \\ 2\alpha + 1 \end{matrix}; \xi\right),$$

up to a constant, and the first boundary condition is satisfied $u(0) = 0$, when $\alpha > 0$.

As is known [28],

$$(14.16) \quad \lim_{\xi \rightarrow 1^-} (1-\xi)^{A+B-C} {}_2F_1\left(\begin{matrix} A, B \\ C \end{matrix}; \xi\right) = \frac{\Gamma(C)\Gamma(A+B-C)}{\Gamma(A)\Gamma(B)},$$

provided $\Re(C - A - B) < 0$. Thus

$$(14.17) \quad \begin{aligned} u(1) &= \lim_{\xi \rightarrow 1^-} (1-\xi) {}_2F_1\left(\begin{matrix} 1 + \alpha + \gamma, 1 + \alpha - \gamma \\ 2\alpha + 1 \end{matrix}; \xi\right) \\ &= \frac{\Gamma(2\alpha + 1)\Gamma(1)}{\Gamma(1 + \alpha + \gamma)\Gamma(1 + \alpha - \gamma)} = 0, \end{aligned}$$

provided that

$$(14.18) \quad \alpha - \gamma = -n = -1, -2, -3, \dots,$$

$$(14.19) \quad \alpha = \alpha_n = \frac{\beta^2 - n^2}{2n} > 0.$$

As a result, the discrete energy levels are given by

$$(14.20) \quad E_n = -V_0 \left(\frac{\beta^2 - n^2}{2\beta n} \right)^2, \quad n = 1, 2, 3, \dots \quad (n^2 < \beta^2).$$

There exists a minimum size of potential hole before any energy eigenvalue at all can be obtained, viz. $\beta^2 = 1$. Equation $1 \leq n^2 \leq \beta^2$ determines the finite number of eigenvalues in a potential hole of a given size [10].

The radial wave functions take the form

$$(14.21) \quad R_n(r) = C_n \xi^\alpha (1 - \xi) {}_2F_1 \left(\begin{matrix} 1 - n, 1 + 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi \right), \quad \xi = e^{-r/a},$$

where the hypergeometric series terminates and C_n is a constant to be determined. Thus, the energy levels can be obtained by the condition (2.7) and the corresponding wave functions are derived with the help of the Rodrigues-type formula (2.8) as follows:

$$(14.22) \quad (\xi^{2\alpha+n-1} (1 - \xi)^n)^{(n-1)} = \frac{\Gamma(2\alpha + n)}{\Gamma(2\alpha + 1)} \xi^{2\alpha} (1 - \xi) {}_2F_1 \left(\begin{matrix} 1 - n, 2\alpha + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi \right).$$

This result follows also, as a special case, from (15.5.9) of [28].

Once again, we can use (13.8) for normalization of the radial wave function.

Then

$$(14.23) \quad aC_n^2 \int_0^1 \xi^{2\alpha-1} (1 - \xi)^2 y_n^2(\xi) d\xi = 1,$$

where

$$(14.24) \quad \xi^{2\alpha} (1 - \xi) y_n(\xi) = \frac{\Gamma(2\alpha + 1)}{\Gamma(2\alpha + n)} [\xi^{2\alpha+n-1} (1 - \xi)^n]^{(n-1)}$$

by (14.22). Moreover,

$$(14.25) \quad (1 - \xi) y_n(\xi) = {}_2F_1 \left(\begin{matrix} -n, 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi \right),$$

by the familiar transformation [27]:

$$(14.26) \quad {}_2F_1 \left(\begin{matrix} A, B \\ C \end{matrix}; \xi \right) = (1 - \xi)^{C-A-B} {}_2F_1 \left(\begin{matrix} C - A, C - B \\ C \end{matrix}; \xi \right)$$

with $A = -n$, $B = 2\alpha + n$, $C = 2\alpha + 1$. Therefore,

$$(14.27) \quad \int_0^1 \xi^{2\alpha-1} (1 - \xi)^2 y_n^2(\xi) d\xi = \frac{\Gamma(2\alpha + 1)}{\Gamma(2\alpha + n)} \\ \times \int_0^1 \left[\xi^{-1} {}_2F_1 \left(\begin{matrix} -n, 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi \right) \right] [\xi^{2\alpha+n-1} (1 - \xi)^n]^{(n-1)} d\xi,$$

and integrating by parts $n - 1$ times, one gets

(14.28)

$$\begin{aligned}
& \int_0^1 \xi^{-1} {}_2F_1\left(\begin{matrix} -n, 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi\right) [\xi^{2\alpha+n-1}(1-\xi)^n]^{(n-1)} d\xi \\
&= \left(\xi^{-1} {}_2F_1\left(\begin{matrix} -n, 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi\right) [\xi^{2\alpha+n-1}(1-\xi)^n]^{(n-2)} \right) \Big|_{\xi=0}^1 \\
&\quad - \int_0^1 \left[\xi^{-1} {}_2F_1\left(\begin{matrix} -n, 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]' [\xi^{2\alpha+n-1}(1-\xi)^n]^{(n-2)} d\xi \\
&= \dots = \\
& (-1)^{k-1} \left(\left[\xi^{-1} {}_2F_1\left(\begin{matrix} -n, 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]^{(k-1)} [\xi^{2\alpha+n-1}(1-\xi)^n]^{(n-k-1)} \right) \Big|_{\xi=0}^1 \\
&\quad + (-1)^k \int_0^1 \left[\xi^{-1} {}_2F_1\left(\begin{matrix} -n, 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]^{(k)} [\xi^{2\alpha+n-1}(1-\xi)^n]^{(n-k-1)} d\xi \\
&= \dots = (-1)^{n-1} \int_0^1 \left[\xi^{-1} {}_2F_1\left(\begin{matrix} -n, 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]^{(n-1)} \xi^{2\alpha+n-1}(1-\xi)^n d\xi,
\end{aligned}$$

in view of the boundary conditions (14.7)–(14.8). By the power series expansion,

(14.29)

$$\begin{aligned}
\left[\xi^{-1} {}_2F_1\left(\begin{matrix} -n, 2\alpha + n \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]^{(n-1)} &= \left(\frac{1}{\xi}\right)^{(n-1)} + \frac{(-n)_n(2\alpha+n)_n}{(n!)(2\alpha+1)_n} (\xi^{n-1})^{(n-1)} \\
&= (-1)^{n-1} \frac{(n-1)!}{\xi^n} + \frac{(-n)_n(2\alpha+n)_n}{(n!)(2\alpha+1)_n} (n-1)!,
\end{aligned}$$

and our integral evaluation can be completed with the aid of the following Euler beta integrals (B.2):

$$\begin{aligned}
\int_0^1 \xi^{2\alpha-1}(1-\xi)^{(n+1)-1} d\xi &= \frac{\Gamma(2\alpha)(n!)}{\Gamma(2\alpha+n+1)}, \\
\int_0^1 \xi^{2\alpha+n-1}(1-\xi)^{(n+1)-1} d\xi &= \frac{\Gamma(2\alpha+n)(n!)}{\Gamma(2\alpha+2n+1)}.
\end{aligned}$$

The final result is given by

$$(14.30) \quad C_n = \frac{(2\alpha)_n}{n!} \sqrt{\frac{(\alpha+n)(2\alpha+n)}{(2\alpha)a}}, \quad \alpha = \frac{\beta^2 - n^2}{2n} \quad (n = 1, 2, 3, \dots)$$

as a complementary normalization in (14.21) (Table 12). We were not able to find the value of this constant in the available literature (see, for example, [10]).

The Hulthén potential at small values of r behaves like a Coulomb potential $U_C = -V_0a/r$, whereas for large values of r it decreases exponentially. (See [10] for more details and a numerical example.) Section 17 below contains an extension of this potential that is suitable for diatomic molecules.

TABLE 12. The Hulthén potential (14.1) in the spherically symmetric case $l = 0$. Substitution: $R(r) = u(\xi)$, $\xi = \exp(-r/a)$, $\alpha^2 = -(2mE/\hbar^2)a^2 > 0$, $\beta^2 = (2mV_0/\hbar^2)a^2 > 0$.

$\sigma(\xi)$	$\xi(1 - \xi)$
$\tilde{\sigma}(\xi)$	$(1 - \xi) [(\alpha^2 + \beta^2)\xi - \alpha^2]$
$\tilde{\tau}(\xi)$	$1 - \xi$
k	$\beta^2 - \alpha$
$\pi(\xi)$	$\alpha - (\alpha + 1)\xi$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$2\alpha + 1 - (2\alpha + 3)\xi$
$\lambda = k + \pi'$	$\beta^2 - 2\alpha - 1 = (\gamma - \alpha - 1)(\gamma + \alpha + 1)$, $\gamma = \sqrt{\alpha^2 + \beta^2}$
$\varphi(\xi)$	$\xi^\alpha(1 - \xi)$
$\rho(\xi)$	$\xi^{2\alpha}(1 - \xi)$
E_n	$-V_0 \left(\frac{\beta^2 - n^2}{2\beta n} \right)^2$, $n = 1, 2, 3, \dots$ ($n^2 < \beta^2$)
$y_n(\xi)$	${}_2F_1(1 - n, 1 + 2\alpha + n; 2\alpha + 1; \xi)$
C_n	$\frac{(2\alpha)_n}{n!} \sqrt{\frac{(\alpha + n)(2\alpha + n)}{(2\alpha)a}}$, $\alpha = \alpha_n = \frac{\beta^2 - n^2}{2n} > 0$

15. Morse Potential

The following central field potential:

$$(15.1) \quad U(r) = D(e^{-2\alpha x} - 2e^{-\alpha x}), \quad x = \frac{r - r_0}{r_0} \quad (0 \leq r < \infty)$$

is used for the study of vibrations of two-atomic molecules [1], [10], [25]. The corresponding Schrödinger equation (5.1) can be solved in spherical coordinates (9.1) when $l = 0$. Introducing new parameters

$$(15.2) \quad \beta^2 = -\frac{2mEr_0^2}{\hbar^2} > 0, \quad \gamma^2 = \frac{2mDr_0^2}{\hbar^2}$$

($\beta, \gamma > 0$), with the help of the following substitution

$$(15.3) \quad R(r) = u(\xi), \quad \xi = \frac{2\gamma}{\alpha} e^{-\alpha x}, \quad x = \frac{r - r_0}{r_0}$$

one gets

$$(15.4) \quad \xi^2 \frac{d^2 u}{d\xi^2} + \xi \frac{du}{d\xi} + \left(-\frac{\beta^2}{\alpha^2} + \frac{\gamma}{\alpha} \xi - \frac{1}{4} \xi^2 \right) u = 0.$$

This is the generalized equation of hypergeometric type with

$$(15.5) \quad \sigma(\xi) = \xi, \quad \tilde{\tau}(\xi) = 1,$$

and

$$(15.6) \quad \tilde{\sigma}(\xi) = -\frac{\beta^2}{\alpha^2} + \frac{\gamma}{\alpha} \xi - \frac{1}{4} \xi^2.$$

The substitution

$$(15.7) \quad u = \varphi(\xi)y(\xi) = \xi^{\beta/\alpha}e^{-\xi/2}y(\xi)$$

results in the confluent hypergeometric equations:

$$(15.8) \quad y'' + \left(\frac{2\beta}{\alpha} + 1 - \xi\right)y' + \left(\frac{\gamma - \beta}{\alpha} - \frac{1}{2}\right)y = 0$$

with the following values of parameters:

$$(15.9) \quad c = 2\frac{\beta}{\alpha} + 1, \quad a = \frac{1}{2}c - \frac{\gamma}{\alpha} = \frac{1}{2} + \frac{\beta - \gamma}{\alpha}$$

(see Table 13 and the corresponding Mathematica file for more details).

The general solution of (15.8) has the form [27], [28]:

$$(15.10) \quad y = A {}_1F_1\left(\begin{matrix} a \\ c \end{matrix}; \xi\right) + B \xi^{1-c} {}_1F_1\left(\begin{matrix} 1 + a - c \\ 2 - c \end{matrix}; \xi\right).$$

Here, the second constant must vanish, $B = 0$, due to the boundary condition $\lim_{r \rightarrow \infty} R(r) = u(0) = 0$, because

$$(15.11) \quad 1 - c + \frac{\beta}{\alpha} = -\frac{\beta}{\alpha} < 0.$$

The first constant A has to be determined by the normalization.

The second boundary condition, namely, $\lim_{r \rightarrow 0} R(r) = u(\xi_0) = 0$ with $\xi_0 = (2\gamma/\alpha)e^\alpha$, states

$$(15.12) \quad {}_1F_1\left(\begin{matrix} a \\ c \end{matrix}; \xi_0\right) = 0,$$

where both coefficients depend on energy in view of (15.2) and (15.9). This transcendental equation for the discrete energy levels cannot be solved explicitly but for all real diatomic molecules $\xi_0 \gg 1$ [10]. This is why one can use the familiar asymptotic:

$$(15.13) \quad {}_1F_1\left(\begin{matrix} a \\ c \end{matrix}; \xi\right) = \frac{\Gamma(c)}{\Gamma(c-a)}(-\xi)^{-a} \left[1 + O\left(\frac{1}{\xi}\right)\right] \\ + \frac{\Gamma(c)}{\Gamma(a)} e^\xi \xi^{a-c} \left[1 + O\left(\frac{1}{\xi}\right)\right], \quad \xi \rightarrow \infty$$

for the confluent hypergeometric function [27], [28].

By eliminating the largest asymptotic term with $\Gamma(a) = \infty$, an approximate quantization rule states:¹

$$(15.14) \quad a = -v; \quad v = 0, 1, 2, \dots; \quad v < \frac{\gamma}{\alpha} - \frac{1}{2}$$

(more details can be found in [10]). The corresponding approximation to the discrete energy levels is given by

$$(15.15) \quad -\beta^2 = -\gamma^2 + 2\gamma\alpha \left(v + \frac{1}{2}\right) - \alpha^2 \left(v + \frac{1}{2}\right)^2.$$

¹ The values $v = 0$ can be added because $\exp(-\xi_0/2) \ll 1$ and the upper bound is due to convergence of the normalization integral (15.19).

This result can also be obtained in the Nikiforov-Uvarov approach by (2.7). Hence, the approximate energy levels in terms of the vibrational quantum number v are

$$(15.16) \quad E_v = -D + \frac{\hbar^2}{2mr_0^2} \left[2\gamma\alpha \left(v + \frac{1}{2} \right) - \alpha^2 \left(v + \frac{1}{2} \right)^2 \right],$$

where the last term reflects the anharmonicity correction. This formula can be rewritten as follows [10]:

$$(15.17) \quad E_v = -D + \hbar\omega \left[\left(v + \frac{1}{2} \right) - \frac{\alpha}{2\gamma} \left(v + \frac{1}{2} \right)^2 \right], \quad \hbar\omega = \hbar^2 \frac{\alpha\gamma}{mr_0^2} = \hbar \frac{\alpha}{r_0} \sqrt{\frac{2D}{m}}.$$

The first two terms in this formula are in complete agreement with the harmonic oscillator energy levels. The last term reflects the anharmonicity correction, which shows that the anharmonic term never exceeds the harmonic one [10].

The corresponding radial wave functions are given in terms of the Laguerre polynomials (Table 19):

$$(15.18) \quad R(r) = R_v(r) = \sqrt{\frac{(2\beta) v!}{r_0 \Gamma(2\beta/\alpha + v + 1)}} \xi^{\beta/\alpha} e^{-\xi/2} L_v^{2\beta/\alpha}(\xi).$$

Here, we have used the following normalization:

$$(15.19) \quad \int_{r=0}^{\infty} R^2(r) dr = \frac{r_0}{\alpha} \int_{\xi=0}^{\xi_0} \xi^{2\beta/\alpha-1} e^{-\xi} y^2(\xi) d\xi \\ \approx C_v^2 \frac{r_0}{\alpha} \int_{\xi=0}^{\infty} \xi^{2\beta/\alpha-1} e^{-\xi} \left(L_v^{2\beta/\alpha}(\xi) \right)^2 d\xi = 1$$

and the following integral:

$$(15.20) \quad I_{-1} = J_{m,m,-1}^{\delta\delta} = \int_0^{\infty} e^{-\xi} \xi^{\delta-1} (L_m^{\delta}(\xi))^2 d\xi = \frac{\Gamma(\delta + m + 1)}{m! \delta}.$$

Here $\delta = 2\beta/\alpha > 0$ (see [41], [42] and Appendix B; further details are left to the reader).

16. Rotation Correction of Morse Potential

The standard centrifugal term [34]:

$$(16.1) \quad \frac{l(l+1)}{r^2} = \frac{l(l+1)}{r_0^2} \frac{1}{(1+x)^2}, \quad x = \frac{r-r_0}{r_0}$$

can be approximated, in the neighborhood of the minimum of the Morse potential $r = r_0$ (or $x = 0$), as follows [10]:

$$(16.2) \quad \frac{l(l+1)}{r^2} \approx \frac{l(l+1)}{r_0^2} (C_0 + C_1 e^{-\alpha x} + C_2 e^{-2\alpha x}),$$

where

$$(16.3) \quad C_0 = 1 - \frac{3}{\alpha} + \frac{3}{\alpha^2}, \quad C_1 = \frac{4}{\alpha} - \frac{6}{\alpha^2}, \quad C_2 = -\frac{1}{\alpha} + \frac{3}{\alpha^2}.$$

Indeed,

$$(16.4) \quad \frac{1}{(1+x)^2} - (C_0 + C_1 e^{-\alpha x} + C_2 e^{-2\alpha x}) = x^3 \left(-\frac{2}{3} \alpha^2 + 3\alpha - 4 \right) + O(x^4), \quad x \rightarrow 0.$$

TABLE 13. The Morse potential (15.1) in the spherically symmetric case $l = 0$. Substitution: $R(r) = u(\xi)$, $\xi = (2\gamma/\alpha) \exp(-\alpha x)$, $\beta^2 = -2mEr_0^2/\hbar^2 > 0$, $\gamma^2 = 2mDr_0^2/\hbar^2$.

$\sigma(\xi)$	ξ
$\tilde{\sigma}(\xi)$	$-\frac{\beta^2}{\alpha^2} + \frac{\gamma}{\alpha}\xi - \frac{1}{4}\xi^2$
$\tilde{\tau}(\xi)$	1
k	$\frac{\gamma - \beta}{\alpha}$
$\pi(\xi)$	$\frac{\beta}{\alpha} - \frac{\xi}{2}$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$1 + \frac{2\beta}{\alpha} - \xi$
$\lambda = k + \pi'$	$\frac{\gamma - \beta}{\alpha} - \frac{1}{2}$
$\varphi(\xi)$	$\xi^{\beta/\alpha} e^{-\xi/2}$
$\rho(\xi)$	$\xi^{2\beta/\alpha} e^{-\xi}$
E_v	$-D + \frac{\hbar^2}{2mr_0^2} \left[2\gamma\alpha \left(v + \frac{1}{2} \right) - \alpha^2 \left(v + \frac{1}{2} \right)^2 \right]$
$y_v(\xi)$	$C_v L_v^{2\beta/\alpha}(\xi) = C_v \frac{\Gamma(2\beta/\alpha + v + 1)}{v! \Gamma(2\beta/\alpha + 1)} {}_1F_1(-v; 2\beta/\alpha + 1; \xi)$
C_v^2	$\frac{(2\beta)v!}{r_0 \Gamma(2\beta/\alpha + v + 1)}$

This consideration allows one to introduce a rotation correction to the Morse potential without changing the mathematical model much [11].

In this approximation, the radial Schrödinger equation (5.1) in spherical coordinates (9.1)

$$(16.5) \quad R''(r) + \left[\frac{2m}{\hbar^2} (E - D (e^{-2\alpha x} - 2e^{-\alpha x})) - \frac{l(l+1)}{r^2} \right] R(r) = 0,$$

with the new variables

$$(16.6) \quad R(r) = u(\xi), \quad \xi = \frac{2\gamma}{\alpha} e^{-\alpha x}, \quad x = \frac{r - r_0}{r_0}$$

and with the modified parameters

$$(16.7) \quad \beta_1^2 = \beta^2 + l(l+1)C_0, \quad \beta^2 = -\frac{2mEr_0^2}{\hbar^2} > 0,$$

$$(16.8) \quad \gamma_1^2 = \gamma^2 - \frac{1}{2}l(l+1)C_1, \quad \gamma_2^2 = \gamma^2 + l(l+1)C_2, \quad \gamma^2 = \frac{2mDr_0^2}{\hbar^2},$$

becomes the following generalized equation of hypergeometric type:

$$(16.9) \quad \xi^2 u'' + \xi u' + \left(-\frac{\beta_1^2}{\alpha^2} + \frac{\gamma_1^2}{\alpha\gamma_2} \xi - \frac{1}{4} \xi^2 \right) u = 0.$$

Here

$$(16.10) \quad \sigma(\xi) = \xi, \quad \tilde{\tau}(\xi) = 1,$$

and

$$(16.11) \quad \tilde{\sigma}(\xi) = -\frac{\beta_1^2}{\alpha^2} + \frac{\gamma_1^2}{\alpha\gamma_2} \xi - \frac{1}{4} \xi^2.$$

The following substitution

$$(16.12) \quad u = \xi^{\beta_1/\alpha} e^{-\xi/2} y(\xi)$$

results, once again, in the confluent hypergeometric equation

$$(16.13) \quad y'' + \left(\frac{2\beta_1}{\alpha} + 1 - \xi \right) y' + \left(\frac{\gamma_1^2}{\alpha\gamma_2} - \frac{\beta_1}{\alpha} - \frac{1}{2} \right) y = 0$$

with the new values of the parameters:

$$(16.14) \quad c_1 = 2\frac{\beta_1}{\alpha} + 1, \quad a_1 = \frac{1}{2}c_1 - \frac{\gamma_1^2}{\alpha\gamma_2} = \frac{1}{2} + \frac{\beta_1}{\alpha} - \frac{\gamma_1^2}{\alpha\gamma_2}$$

(see Table 14 and the corresponding Mathematica file for more details).

An approximate quantization rule states:

$$(16.15) \quad a_1 = -v; \quad v = 0, 1, 2, \dots$$

Therefore,

$$(16.16) \quad -\beta_1^2 = -\left[\frac{\gamma_1^2}{\gamma_2} - \alpha \left(v + \frac{1}{2} \right) \right]^2,$$

and, in the energy formula, one has to replace γ by

$$(16.17) \quad \frac{\gamma_1^2}{\gamma_2} \approx \gamma \left[1 - l(l+1) \frac{C_1 + C_2}{2\gamma^2} \right], \quad \gamma \gg 1.$$

As a result, we arrive at the following vibration-rotation energy levels:

$$(16.18) \quad E = E_{vl} = \frac{\hbar^2}{2mr_0^2} \left[-\gamma^2 + 2\gamma\alpha \left(v + \frac{1}{2} \right) - \alpha^2 \left(v + \frac{1}{2} \right)^2 + l(l+1) - \frac{3(\alpha-1)}{\alpha\gamma} \left(v + \frac{1}{2} \right) l(l+1) - \frac{9(\alpha-1)^2}{4\alpha^4\gamma^2} l^2(l+1)^2 \right].$$

This formula can be presented in the form

$$(16.19) \quad E_{vl} = -D + \hbar\omega \left[\left(v + \frac{1}{2} \right) - \frac{\alpha}{2\gamma} \left(v + \frac{1}{2} \right)^2 \right] + \frac{\hbar^2 l(l+1)}{2mr_0^2} - \frac{3(\alpha-1)}{2\alpha^2 D} \hbar\omega \left(v + \frac{1}{2} \right) \frac{\hbar^2 l(l+1)}{2mr_0^2} - \frac{9(\alpha-1)^2}{4\alpha^2 D} \left(\frac{\hbar^2 l(l+1)}{2mr_0^2} \right)^2,$$

where

$$(16.20) \quad \hbar\omega = \frac{\hbar^2 \alpha \gamma}{mr_0^2} = \hbar \left(\frac{\alpha}{r_0} \sqrt{\frac{2D}{m}} \right).$$

TABLE 14. The rotation correction of the Morse potential (16.2), where the coefficients C_0 , C_1 , and C_2 are given by (16.3).

$\sigma(\xi)$	ξ
$\tilde{\sigma}(\xi)$	$-\frac{\beta_1^2}{\alpha^2} + \frac{\gamma_1^2}{\alpha\gamma_2}\xi - \frac{1}{4}\xi^2$
$\tilde{\tau}(\xi)$	1
k	$\frac{\gamma_1^2}{\alpha\gamma_2} - \frac{\beta_1}{\alpha}$
$\pi(\xi)$	$\frac{\beta_1}{\alpha} - \frac{\xi}{2}$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$1 + \frac{2\beta_1}{\alpha} - \xi$
$\lambda = k + \pi'$	$\frac{\gamma_1^2}{\alpha\gamma_2} - \frac{\beta_1}{\alpha} - \frac{1}{2}$
$\varphi(\xi)$	$\xi^{\beta_1/\alpha} e^{-\xi/2}$
$\rho(\xi)$	$\xi^{2\beta_1/\alpha} e^{-\xi}$

The first three terms of this formula are exactly the same as those derived in the previous case; see (15.17). The fourth term can be interpreted as the molecule rotational energy at fixed distance r_0 . The next term represents a coupling of the vibrations and rotations, which is negative because at higher vibrational quantum numbers the average nuclear distance increases beyond r_0 in consequence of the anharmonicity [10]. The last term can be thought of as a negative second-order correction to the rotation energy.

The corresponding radial wave functions are given in terms of the Laguerre polynomials (Table 19):

$$(16.21) \quad R(r) = R_{vl}(r) = \sqrt{\frac{(2\beta_1) v!}{r_0 \Gamma(2\beta_1/\alpha + v + 1)}} \xi^{\beta_1/\alpha} e^{-\xi/2} L_v^{2\beta_1/\alpha}(\xi).$$

Once again, we have used the normalization similar to (15.19). Further details on the rotational corrections of Morse formulas are discussed in [10], [11].

17. Modified Hulthén Potential

Once again, we are looking for solutions of the stationary Schrödinger equation (5.1) in spherical coordinates (9.1) for the following central field potential,

$$(17.1) \quad U(r) = -V_0 \frac{e^{-r/a} (1 - be^{-r/a})}{(1 - e^{-r/a})^2} \quad (0 \leq r < \infty),$$

when $b > 1$ and $l = 0$. (The special case $b = 1$ corresponds to the original Hulthén potential (14.1) above.) By using substitution (14.2) with the same positive parameters (14.3), one obtains the following generalized equation of hypergeometric

type,

$$(17.2) \quad \xi^2 \frac{d^2 u}{d\xi^2} + \xi \frac{du}{d\xi} + \left[-\alpha^2 + \beta^2 \frac{\xi(1-b\xi)}{(1-\xi)^2} \right] u = 0,$$

where

$$(17.3) \quad \sigma(\xi) = \xi(1-\xi), \quad \tilde{\tau}(\xi) = 1-\xi,$$

and

$$(17.4) \quad \tilde{\sigma}(\xi) = -(\alpha^2 + b\beta^2)\xi^2 + (2\alpha^2 + \beta^2)\xi - \alpha^2.$$

For the modified Hulthén potential, let us first analyze the same boundary conditions (14.8) as before. We have

$$(17.5) \quad k = \beta^2 \mp 2\alpha\kappa, \quad \pi(\xi) = -\frac{\xi}{2} \pm [(\alpha \pm \kappa)\xi - \alpha],$$

where

$$(17.6) \quad \kappa = \sqrt{\frac{1}{4} + (b-1)\beta^2} > \frac{1}{2},$$

for all four possible sign combinations. The solutions can be found by putting:

$$(17.7) \quad u = \varphi(\xi)y(\xi) = \xi^\alpha(1-\xi)^{\kappa+1/2}y(\xi),$$

which results in the hypergeometric equation (14.13):

$$(17.8) \quad \xi(1-\xi)y'' + [2\alpha+1-2(\alpha+\kappa+1)\xi]y' - \left[(2\alpha+1)\left(\kappa+\frac{1}{2}\right) - \beta^2 \right] y = 0,$$

with

$$(17.9) \quad A = \alpha + \kappa + \frac{1}{2} + \gamma, \quad B = \alpha + \kappa + \frac{1}{2} - \gamma, \quad C = 2\alpha + 1, \quad \gamma = \sqrt{\alpha^2 + b\beta^2}$$

(details of calculations are presented in Table 16 and in a complementary Mathematica file).

The required solution, that is bounded at $\xi = 0$, has the form

$$(17.10) \quad y = {}_2F_1\left(\begin{matrix} 1/2 + \kappa + \alpha + \gamma, & 1/2 + \kappa + \alpha - \gamma \\ & 2\alpha + 1 \end{matrix}; \xi\right),$$

up to a constant, and the first boundary condition is satisfied $u(0) = 0$, when $\alpha > 0$.

In view of (14.16), one gets

$$(17.11) \quad u(1) = \lim_{\xi \rightarrow 1^-} (1-\xi)^{-(\kappa-1/2)} \left[(1-\xi)^{2\kappa} {}_2F_1\left(\begin{matrix} 1/2 + \kappa + \alpha + \gamma, & 1/2 + \kappa + \alpha - \gamma \\ & 2\alpha + 1 \end{matrix}; \xi\right) \right]$$

$$= \frac{\Gamma(2\alpha+1)\Gamma(2\kappa)}{\Gamma(1/2+\kappa+\alpha+\gamma)\Gamma(1/2+\kappa+\alpha-\gamma)} \times \lim_{\xi \rightarrow 1^-} (1-\xi)^{-(\kappa-1/2)} = \infty.$$

Therefore, one has to look for square-integrable solutions that are not bounded at the origin.

By terminating the hypergeometric series in (17.10), or by using the Nikiforov-Uvarov condition (2.7), we arrive at the following quantization rule:

$$(17.12) \quad \alpha + \kappa + \frac{1}{2} - \gamma = -n \quad (n = 0, 1, 2, \dots \leq n_{\max}),$$

or

$$(17.13) \quad \alpha = \frac{b\beta^2 - (n + \kappa + 1/2)^2}{2(n + \kappa + 1/2)} > 0.$$

The corresponding energy levels are given by

$$(17.14) \quad E_n = -V_0 \left(\frac{b\beta^2 - (n + \kappa + 1/2)^2}{2\beta(n + \kappa + 1/2)} \right)^2,$$

where

$$(17.15) \quad b\beta^2 - \left(\kappa + n + \frac{1}{2} \right)^2 = \beta^2 - n^2 - (2\kappa + 1) \left(n + \frac{1}{2} \right) > 0.$$

Once again, there exists a minimum size of potential hole before any energy eigenvalue at all can be obtained, viz. $b\beta^2 = (\kappa + 1/2)^2$. Equation $(n + \kappa + 1/2)^2 \leq b\beta^2$ determines the finite number of eigenvalues, $n \leq n_{\max}$, in a potential hole of a given size.

The radial wave functions take the form

$$(17.16) \quad R_n(r) = C_n \xi^\alpha (1 - \xi)^{\kappa+1/2} {}_2F_1 \left(\begin{matrix} -n, 2\alpha + 2\kappa + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi \right), \quad \xi = e^{-r/a},$$

where the hypergeometric series terminates and C_n is a constant to be determined. Once again, this result can also be obtained with the help of the Rodrigues-type formula (2.8):

$$(17.17) \quad [\xi^{2\alpha+n} (1 - \xi)^{2\kappa+n}]^{(n)} = \frac{\Gamma(2\alpha + n + 1)}{\Gamma(2\alpha + 1)} \xi^{2\alpha} (1 - \xi)^{2\kappa} {}_2F_1 \left(\begin{matrix} -n, 2\alpha + 2\kappa + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi \right).$$

(It can be thought of as a special case of (15.5.9) from [28].)

The normalization condition (13.8) for the radial wave function becomes

$$(17.18) \quad aC_n^2 \int_0^1 \xi^{2\alpha-1} (1 - \xi)^{2\kappa+1} y_n^2(\xi) d\xi = 1,$$

where

$$(17.19) \quad \xi^{2\alpha} (1 - \xi)^{2\kappa} y_n(\xi) = \frac{\Gamma(2\alpha + 1)}{\Gamma(2\alpha + n + 1)} [\xi^{2\alpha+n} (1 - \xi)^{2\kappa+n}]^{(n)}.$$

Integrating by parts n times, one gets

(17.20)

$$\begin{aligned}
& \int_0^1 (\xi^{-1} - 1) {}_2F_1\left(\begin{matrix} -n, 2\alpha + 2\kappa + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi\right) [\xi^{2\alpha+n}(1-\xi)^{2\kappa+n}]^{(n)} d\xi \\
&= \left((\xi^{-1} - 1) {}_2F_1\left(\begin{matrix} -n, 2\alpha + 2\kappa + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi\right) [\xi^{2\alpha+n}(1-\xi)^{2\kappa+n}]^{(n-1)} \right) \Big|_{\xi=0}^1 \\
&\quad - \int_0^1 \left[(\xi^{-1} - 1) {}_2F_1\left(\begin{matrix} -n, 2\alpha + 2\kappa + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]' [\xi^{2\alpha+n}(1-\xi)^{2\kappa+n}]^{(n-1)} d\xi \\
&= \dots = \\
& (-1)^{k-1} \left(\left[(\xi^{-1} - 1) {}_2F_1\left(\begin{matrix} -n, 2\alpha + 2\kappa + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]^{(k-1)} [\xi^{2\alpha+n}(1-\xi)^{2\kappa+n}]^{(n-k)} \right) \Big|_{\xi=0}^1 \\
&+ (-1)^k \int_0^1 \left[(\xi^{-1} - 1) {}_2F_1\left(\begin{matrix} -n, 2\alpha + 2\kappa + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]^{(k)} [\xi^{2\alpha+n}(1-\xi)^{2\kappa+n}]^{(n-k)} d\xi \\
&= \dots = (-1)^n \int_0^1 \left[(\xi^{-1} - 1) {}_2F_1\left(\begin{matrix} -n, 2\alpha + 2\kappa + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]^{(n)} \xi^{2\alpha+n}(1-\xi)^{2\kappa+n} d\xi.
\end{aligned}$$

By the power series expansion,

$$\begin{aligned}
& \left[(\xi^{-1} - 1) {}_2F_1\left(\begin{matrix} -n, 2\alpha + 2\kappa + n + 1 \\ 2\alpha + 1 \end{matrix}; \xi\right) \right]^{(n)} \\
&= \left(\frac{1}{\xi} \right)^{(n)} - \frac{(-n)_n (2\alpha + 2\kappa + n + 1)_n}{(n!)(2\alpha + 1)_n} (\xi^n)^{(n)} \\
&= (-1)^n \frac{(n!)}{\xi^{n+1}} - \frac{(-n)_n (2\alpha + 2\kappa + n + 1)_n}{(n!)(2\alpha + 1)_n} (n!),
\end{aligned}$$

and the integral evaluation can be completed with the aid of the following Euler beta integrals (B.2):

$$\begin{aligned}
\int_0^1 \xi^{2\alpha-1} (1-\xi)^{(2\kappa+n+1)-1} d\xi &= \frac{\Gamma(2\alpha)\Gamma(2\kappa+n+1)}{\Gamma(2\alpha+2\kappa+n+1)}, \\
\int_0^1 \xi^{2\alpha+n} (1-\xi)^{2\kappa+n} d\xi &= \frac{\Gamma(2\alpha+n+1)\Gamma(2\kappa+n+1)}{\Gamma(2\alpha+2\kappa+2n+2)}.
\end{aligned}$$

The final result is given by

$$(17.21) \quad C_n^2 = \frac{\Gamma(2\alpha+n+1)\Gamma(2\alpha+2\kappa+n+1)(2\alpha+2\kappa+2n+1)}{a(2\alpha)\Gamma^2(2\alpha)\Gamma(2\kappa+n+1)(2\kappa+2n+1)n!}.$$

Our modification of the Hulthén potential can be used for study of vibrations for diatomic molecules, when m becomes the reduced mass of two atoms. Therefore, it is informative to compare the classical Morse potential and the modified Hulthén one. Suppose that at the common point of the potential minimum $r = r_{\min}$ the

TABLE 15. Comparison of the parameters for the Morse potential and the modified Hulthén potential: $E(\text{eV}) = E(\text{cm}^{-1}) \times 1.2398 \times 10^{-4}$ according to [10].

Molecule	$\frac{\hbar^2}{2mr_0^2}$ (cm ⁻¹)	D (cm ⁻¹)	α	b	V_0 (cm ⁻¹)
H ₂	60.8296	38292	1.440	1.5904	67394
HCl	10.5930	37244	2.380	4.51744	524010
I ₂	0.0374	12550	4.954	68.848	198490

following conditions hold:

$$(17.22) \quad \begin{aligned} U_{\text{Morse}}(r_{\min}) &= U_{\text{ModHulthén}}(r_{\min}), \\ \frac{d}{dr}U_{\text{Morse}}(r_{\min}) &= \frac{d}{dr}U_{\text{ModHulthén}}(r_{\min}) = 0, \\ \frac{d^2}{dr^2}U_{\text{Morse}}(r_{\min}) &= \frac{d^2}{dr^2}U_{\text{ModHulthén}}(r_{\min}). \end{aligned}$$

Then $r_{\min} = r_0$, $\exp(r_0/a) = 2b - 1$ and

$$(17.23) \quad D = \frac{V_0}{4(b-1)}, \quad 2D \frac{\alpha^2}{r_0^2} = \frac{(2b-1)^2 V_0}{8a^2(b-1)^3}$$

(see our Mathematica file for more details). The Morse potential (15.1) can be rewritten in an equivalent form:

$$(17.24) \quad \begin{aligned} U_{\text{Morse}}(r) &= \frac{V_0}{4(b-1)} (2b-1)^{\frac{2b-1}{2(b-1)}} \\ &\times \left[(2b-1)^{\frac{2b-1}{2(b-1)}} \exp\left(-\frac{(2b-1)r}{(b-1)a}\right) - 2 \exp\left(-\frac{(2b-1)r}{2(b-1)a}\right) \right], \end{aligned}$$

where the parameters of the modified Hulthén potential (17.1) have been utilized.

A graphical example, comparing both potentials, when $V_0 = a = 1$, $b = 2$, and $r_{\min} = \ln 3 \simeq 1.09861$, is presented in Figure 2 and further discussed in our Mathematica notebook for the reader's convenience, as well as potentials for molecules H₂, HCl, and I₂ with a completion of the data from [10] in Table 15 (see also Figure 3).

In order to find parameters of the modified Hulthén potential in terms of the Morse ones, we have solved numerically the following equation:

$$(17.25) \quad \alpha = \varphi(b) := \frac{(2b-1) \ln(2b-1)}{2(b-1)}, \quad b \geq 1$$

with the monotone function $\varphi(b)$. Then $V_0 = 4(b-1)D$ and $a = r_0/\ln(2b-1)$. (More details are provided in the Mathematica file.)

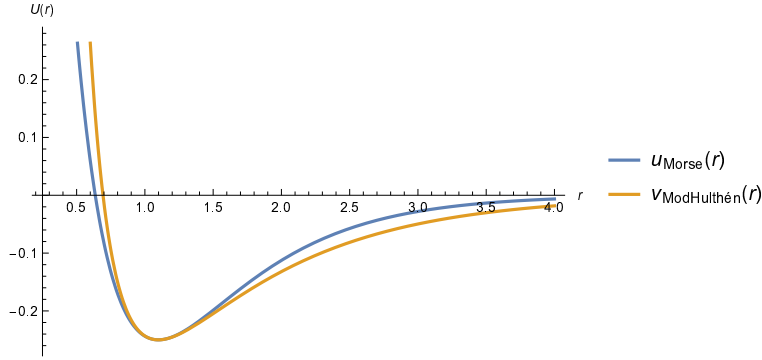


FIGURE 2. An example, comparing the Morse potential (blue) and the modified Hulthén potential (brown), when $V_0 = a = 1$, $b = 2$, and $r_{\min} = \ln 3 \simeq 1.09861$.

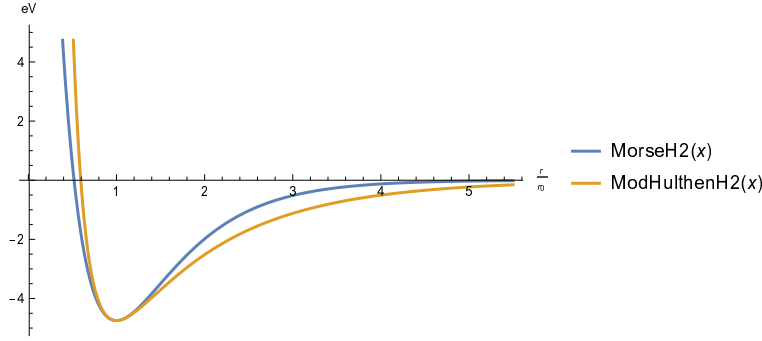


FIGURE 3. The Morse potential (blue) [10] and the modified Hulthén potential (brown), for the molecule H_2 .

Remark. The so-called generalized Morse potential is usually introduced as

$$(17.26) \quad U(r) = D \left(1 - \frac{e^{\alpha r_0} - 1}{e^{\alpha r} - 1} \right)^2$$

(D, α, r_0 are parameters) with the following asymptotic

$$(17.27) \quad U(\infty) = \lim_{r \rightarrow \infty} U(r) = D > 0$$

(see [9], [31], and the references therein). The difference

$$(17.28) \quad U(r) - U(\infty) = -2D(e^{\alpha r_0} - 1) \frac{1 - e^{-\alpha r}(1 + e^{\alpha r_0})/2}{(1 - e^{-\alpha r})^2} e^{-\alpha r}$$

has the form of our modified Hulthén potential (17.1) with the parameters given by

$$(17.29) \quad V_0 = 2D(e^{\alpha r_0} - 1), \quad b = \frac{1}{2}(1 + e^{\alpha r_0}), \quad a = \frac{1}{\alpha}.$$

Therefore, this case has been studied here as well (see also section 19 for an independent computer algebra approach).

TABLE 16. The modified Hulthén potential (17.1) in the spherically symmetric case $l = 0$. Substitution: $R(r) = u(\xi)$, $\xi = \exp(-r/a)$, $\alpha^2 = -(2mE/\hbar^2) a^2 > 0$, $\beta^2 = (2mV_0/\hbar^2) a^2 > 0$.

$\sigma(\xi)$	$\xi(1 - \xi)$
$\tilde{\sigma}(\xi)$	$-(\alpha^2 + b\beta^2)\xi^2 + (2\alpha^2 + \beta^2)\xi - \alpha^2$
$\tilde{\tau}(\xi)$	$1 - \xi$
k	$\beta^2 - 2\alpha\kappa$, $\kappa = \sqrt{(1/4) + (b-1)\beta^2}$
$\pi(\xi)$	$\alpha - (\alpha + \kappa + 1/2)\xi$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$2\alpha + 1 - 2(\alpha + \kappa + 1)\xi$
$\lambda = k + \pi'$	$\beta^2 - (2\alpha + 1)(\kappa + 1/2) = (\gamma - \alpha - \kappa - 1/2)(\gamma + \alpha + \kappa + 1/2)$
$\varphi(\xi)$	$\xi^\alpha(1 - \xi)^{\kappa+1/2}$, $\gamma = \sqrt{\alpha^2 + b\beta^2}$
$\rho(\xi)$	$\xi^{2\alpha}(1 - \xi)^{2\kappa}$
E_n	$-V_0 \left(\frac{b\beta^2 - (n + \kappa + 1/2)^2}{2\beta(n + \kappa + 1/2)} \right)^2$, $n = 0, 1, 2, \dots$ ($n \leq n_{\max}$)
$y_n(\xi)$	${}_2F_1(-n, 2\alpha + 2\kappa + n + 1; 2\alpha + 1; \xi)$, $\alpha = \alpha_n = \frac{b\beta^2 - (n + \kappa + 1/2)^2}{2(n + \kappa + 1/2)} > 0$
C_n^2	$\frac{\Gamma(2\alpha + n + 1)\Gamma(2\alpha + 2\kappa + n + 1)(2\alpha + 2\kappa + 2n + 1)}{a(2\alpha)\Gamma^2(2\alpha)\Gamma(2\kappa + n + 1)(2\kappa + 2n + 1)n!}$

18. Rotation Correction of Modified Hulthén Potential

In view of

$$(18.1) \quad \frac{dU(r)}{dr} = V_0 e^{r/a} \frac{1 - 2b + e^{r/a}}{a(e^{r/a} - 1)^3} = 0,$$

the minimum of the modified Hulthén potential (17.1) occurs when

$$(18.2) \quad e^{r_{\min}/a} = 2b - 1 > 1.$$

By letting $r = a(1+x)x_0$ in the neighborhood of this minimum $r \approx r_{\min}$ (or $x \approx 0$), with the aid of Mathematica, we derive the following expansion:

$$(18.3) \quad \frac{1}{(1+x)^2} - C_0 - e^{-(1+x)x_0} \frac{C_1 + C_2 e^{-(1+x)x_0}}{(1 - e^{-(1+x)x_0})^2} \\ = \left(-4 + \frac{3bx_0}{b-1} - \frac{1-2b+4b^2}{6(b-1)^2} x_0^2 \right) x^3 + O(x^4), \quad x \rightarrow 0,$$

where

$$(18.4) \quad \begin{aligned} C_0 &= 1 + 4(b-1) \frac{3(b-1) - (3b-1)x_0}{(2b-1)^2 x_0^2}, \\ C_1 &= 8(b-1)^2 \frac{6(1-b) + (4b-1)x_0}{(2b-1)^2 x_0^2}, \\ C_2 &= 8(b-1)^2 \frac{6b(b-1) + (1-2b(b+1))x_0}{(2b-1)^2 x_0^2} \end{aligned}$$

and $e^{x_0} = 2b-1 > 1$ (see our complementary Mathematica file).

Therefore,

$$(18.5) \quad \frac{l(l+1)}{r^2} \approx \frac{l(l+1)}{r_{\min}^2} \left[C_0 + e^{-r/a} \frac{C_1 + C_2 e^{-r/a}}{(1 - e^{-r/a})^2} \right] \quad (r_{\min} = ax_0),$$

and our equation (17.2) for the modified Hulthén potential holds, once again, but with the following modified values of the parameters:

$$(18.6) \quad \begin{aligned} \alpha^2 \rightarrow \alpha_1^2 &= \alpha^2 + \frac{l(l+1)}{x_0^2} C_0, & \beta^2 \rightarrow \beta_1^2 &= \beta^2 - \frac{l(l+1)}{x_0^2} C_1, \\ b \rightarrow b_1 &= \frac{\beta_2^2}{\beta_1^2}, & \beta_2^2 &= b\beta^2 + \frac{l(l+1)}{x_0^2} C_2, \end{aligned}$$

namely,

$$(18.7) \quad \xi^2 \frac{d^2 u}{d\xi^2} + \xi \frac{du}{d\xi} + \left[-\alpha_1^2 + \beta_1^2 \frac{\xi(1-b_1\xi)}{(1-\xi)^2} \right] u = 0.$$

As a result, we have arrived at the generalized equation of hypergeometric type (Table 17) and our analysis from the previous section is valid, say, up to a proper change of parameters.

For example, the vibration and rotation energy levels are given by

$$(18.8) \quad \begin{aligned} E_{v,l} &= -V_0 \left(\frac{\beta_2^2 - (\kappa_1 + v + 1/2)^2}{2\beta(\kappa_1 + v + 1/2)} \right)^2 \\ &\quad + \frac{\hbar^2 l(l+1)}{2mr_{\min}^2} \left[1 + 4(b-1) \frac{3(b-1) - (3b-1)x_0}{(2b-1)^2 x_0^2} \right] \end{aligned}$$

($v = 0, 1, 2, \dots \leq v_{\max}$). Here,

$$(18.9) \quad \beta_2^2 - \left(\kappa_1 + v + \frac{1}{2} \right)^2 = \beta_1^2 - v^2 - (2\kappa_1 + 1) \left(v + \frac{1}{2} \right) > 0.$$

The corresponding normalized wave functions are the same as before in (17.16) and (17.21) but with the new values of parameters (18.6) (we leave further details to the reader).

Our study of the rotation correction of the modified Hulthén potential will be continued elsewhere (see also the Mathematica file).

TABLE 17. The rotation correction of the modified Hulthén potential (18.5), where the coefficients C_0 , C_1 , and C_2 are given by (18.4).

$\sigma(\xi)$	$\xi(1 - \xi)$
$\tilde{\sigma}(\xi)$	$-(\alpha_1^2 + b_1\beta_1^2)\xi^2 + (2\alpha_1^2 + \beta_1^2)\xi - \alpha_1^2$
$\tilde{\tau}(\xi)$	$1 - \xi$
k	$\beta_1^2 - 2\alpha_1\kappa_1, \quad \kappa_1 = \sqrt{(1/4) + (b_1 - 1)\beta_1^2}$
$\pi(\xi)$	$\alpha_1 - (\alpha_1 + \kappa_1 + 1/2)\xi$
$\tau(\xi) = \tilde{\tau} + 2\pi$	$(2\alpha_1 + 1) - 2(\alpha_1 + \kappa_1 + 1)\xi$
$\lambda = k + \pi'$	$\beta_1^2 - (2\alpha_1 + 1)(\kappa_1 + 1/2)$
$\varphi(\xi)$	$\xi^{\alpha_1}(1 - \xi)^{\kappa_1 + 1/2}$
$\rho(\xi)$	$\xi^{2\alpha_1}(1 - \xi)^{2\kappa_1}$

19. Generalized Morse Potential

The potential of the form:

$$(19.1) \quad U(r) = D \left(1 - \frac{\gamma}{e^{ar} - 1} \right)^2, \quad \gamma = e^{ar_0} - 1 \quad (0 \leq r < \infty)$$

by (17.28) is reduced to the modified Hulthén potential (17.1) up to a proper change of parameters (17.29). This is why, we can use solutions from section 17.

On the second thought, the standard change of the variables [9]:

$$(19.2) \quad R(r) = \sqrt{a}u(\eta), \quad \eta = (e^{ar} - 1)^{-1}$$

results in the generalized equation of hypergeometric type with the following coefficients:

$$(19.3) \quad \sigma(\eta) = \eta(1 + \eta), \quad \tilde{\sigma}(\eta) = \epsilon - \kappa(1 - \gamma\eta)^2, \quad \tilde{\tau}(\eta) = 2\eta + 1,$$

where

$$(19.4) \quad \kappa = \frac{2mD}{a^2\hbar^2}, \quad \epsilon = \frac{2mE}{a^2\hbar^2}; \quad \alpha^2 = \kappa - \epsilon > 0, \quad \beta^2 = \kappa(\gamma + 1)^2 - \epsilon > 0$$

(see Table 18 and the Mathematica file). The following substitution

$$(19.5) \quad u = \varphi(\eta)y(\eta) = \frac{\eta^\alpha}{(1 + \eta)^\beta}y(\eta)$$

results in the hypergeometric equation of Gauss (14.13) in the variable $-\eta$ with the following values of parameters:

$$(19.6) \quad A = \alpha - \beta + \delta, \quad B = \alpha - \beta - \delta + 1, \quad C = 2\alpha + 1; \quad \delta = \frac{1}{2} + \sqrt{\frac{1}{4} + \kappa\gamma^2}.$$

Bound states correspond to the polynomial solutions, when $A = \alpha - \beta + \delta = -n$ and n are some nonnegative integers. Thus α and β are n -dependent, once again, and satisfy the following equations:

$$(19.7) \quad \beta_n - \alpha_n = n + \delta, \quad \beta_n^2 - \alpha_n^2 = \kappa\gamma(\gamma + 2).$$

TABLE 18. The generalized Morse potential (19.1). The coefficients ϵ and κ are given by (19.4).

$\sigma(\eta)$	$\eta(1 + \eta)$
$\tilde{\sigma}(\eta)$	$\epsilon - \kappa(1 - \gamma\eta)^2, \quad \kappa\gamma^2 = \delta(\delta - 1)$
$\tilde{\tau}(\eta)$	$2\eta + 1$
k	$2(\alpha^2 - \alpha\beta + \kappa\gamma) = (\alpha - \beta)^2 - \delta(\delta - 1)$
$\pi(\eta)$	$\alpha + (\alpha - \beta)\eta, \quad \alpha = \sqrt{\kappa - \epsilon}, \quad \beta = \sqrt{\kappa(\gamma + 1)^2 - \epsilon}$
$\tau(\eta) = \tilde{\tau} + 2\pi$	$2\alpha + 1 + 2(\alpha - \beta + 1)\eta$
$\lambda = k + \pi'$	$\alpha - \beta + 2\alpha^2 - 2\alpha\beta + 2\kappa\gamma = (\alpha - \beta + \delta)(\alpha - \beta - \delta + 1)$
$\varphi(\eta)$	$\eta^\alpha(1 + \eta)^{-\beta}$
$\rho(\eta)$	$\eta^{2\alpha}(1 + \eta)^{-2\beta}$

One finds that

$$(19.8) \quad \alpha_n = \frac{1}{2} \left(\frac{\kappa\gamma(\gamma + 2)}{n + \delta} - n - \delta \right), \quad \beta_n = \frac{1}{2} \left(\frac{\kappa\gamma(\gamma + 2)}{n + \delta} + n + \delta \right)$$

and the energy levels have the form [9]:

$$(19.9) \quad E_n = D - \frac{\alpha^2 \hbar^2}{8m} \left(n + \delta - \frac{\kappa\gamma(\gamma + 2)}{n + \delta} \right)^2.$$

The corresponding normalized wave functions are given by

$$(19.10) \quad R_n(r) = \sqrt{a} C_n e^{-\beta_n ar} (e^{ar} - 1)^{\beta_n - \alpha_n} {}_2F_1 \left(\begin{matrix} -n, 1 - n - 2\delta \\ 2\alpha_n + 1 \end{matrix}; \frac{1}{1 - e^{ar}} \right) \\ = \sqrt{a} C_n e^{-\alpha_n ar} (1 - e^{-ar})^\delta {}_2F_1 \left(\begin{matrix} -n, 2\alpha_n + 2\delta + n \\ 2\alpha_n + 1 \end{matrix}; e^{-ar} \right)$$

[we have used (15.8.1) of [28]] in order to obtain the standard form (17.16)], with the normalization coefficients found in [9] as follows

$$(19.11) \quad C_n^2 = \frac{a(\alpha_n + n + \delta)\Gamma(2\alpha_n + n + 1)\Gamma(2\alpha_n + n + 2\delta)}{n!(n + \delta)\Gamma(2\alpha_n)\Gamma(2\alpha_n + 1)\Gamma(n + 2\delta)},$$

due to the normalization condition:

$$(19.12) \quad \int_0^\infty R_n^2(r) dr = 1.$$

There is only a finite set of the energy levels:

$$(19.13) \quad n = 0, 1, 2, \dots, n_{\max} \leq \sqrt{\kappa\gamma(\gamma + 2)} - \delta.$$

One can use our results from section 17 in order to verify all these formulas, originally presented in [9] (see also [1], [31], and the references therein). We leave further details to the reader.

In a similar fashion, one can consider the Wood–Saxon potential and the motion of the electron in magnetic field [1], [9], [10], [15] [23], and [27].

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A. Data for the Classical Orthogonal Polynomials

The basic information about classical orthogonal polynomials, namely, for the Jacobi $P_n^{(\alpha,\beta)}(x)$, Laguerre $L_n^\alpha(x)$, and Hermite $H_n(x)$ polynomials, is presented, for the reader's convenience, in Table 19. It contains the coefficients of the differential equation (2.3), the intervals of orthogonality (a, b) , the weight functions $\rho(x)$ and constants B_n in the Rodrigues-type formula (2.8), the leading terms:

$$(A.1) \quad y_n(x) = a_n x^n + b_n x^{n-1} + \dots$$

for these polynomials, their squared norms:

$$(A.2) \quad d_n^2 = \int_a^b y_n^2(x) \rho(x) dx,$$

and the coefficients of the three-term recurrence relation:

$$(A.3) \quad x y_n(x) = \alpha_n y_{n+1}(x) + \beta_n y_n(x) + \gamma_n y_{n-1}(x),$$

where

$$(A.4) \quad \alpha_n = \frac{a_n}{a_{n+1}}, \quad \beta_n = \frac{b_n}{a_n} - \frac{b_{n+1}}{a_{n+1}}, \quad \gamma_n = \alpha_{n-1} \frac{d_n^2}{d_{n-1}^2}.$$

(More details can be found in [2], [26], [27], [28], and [42].)

B. An Integral Evaluation

The following useful integral:

$$(B.1) \quad \begin{aligned} J_{nms}^{\alpha\beta} &= \int_0^\infty e^{-x} x^{\alpha+s} L_n^\alpha(x) L_m^\beta(x) dx \\ &= (-1)^{n-m} \frac{\Gamma(\alpha+s+1) \Gamma(\beta+m+1) \Gamma(s+1)}{m! (n-m)! \Gamma(\beta+1) \Gamma(s-n+m+1)} \\ &\quad \times {}_3F_2\left(\begin{matrix} -m, s+1, \beta-\alpha-s \\ \beta+1, n-m+1 \end{matrix}; 1\right), \quad n \geq m, \end{aligned}$$

where parameter s may take some integer values and ${}_3F_2(1)$ is the generalized hypergeometric series [2], [28], has been evaluated in [41] and [42] (see also [13], [35]). Special cases have been used above for the normalization of the wave functions; see (7.13) and (15.20).

We use also throughout the manuscript the familiar Euler beta and gamma integrals:

$$(B.2) \quad B(\alpha, \beta) = \int_0^1 t^{\alpha-1} (1-t)^{\beta-1} dt = \frac{\Gamma(\alpha)\Gamma(\beta)}{\Gamma(\alpha+\beta)},$$

provided $\Re(\alpha) > 0$ and $\Re(\beta) > 0$,

$$(B.3) \quad \Gamma(\alpha) = \int_0^\infty t^{\alpha-1} e^{-t} dt, \quad \Re(\alpha) > 0.$$

(See [2], [27], and [28] for more details.)

TABLE 19. Data for the Jacobi $P_n^{(\alpha,\beta)}(x)$, Laguerre $L_n^\alpha(x)$, and Hermite $H_n(x)$ polynomials.

$y_n(x)$	$P_n^{(\alpha,\beta)}(x) (\alpha > -1, \beta > -1)$	$L_n^\alpha(x) (\alpha > -1)$	$H_n(x)$
(a, b)	$(-1, 1)$	$(0, \infty)$	$(-\infty, \infty)$
$\rho(x)$	$(1-x)^\alpha(1+x)^\beta$	$x^\alpha e^{-x}$	e^{-x^2}
$\sigma(x)$	$1-x^2$	x	1
$\tau(x)$	$\beta - \alpha - (\alpha + \beta + 2)x$	$1 + \alpha - x$	$-2x$
λ_n	$n(\alpha + \beta + n + 1)$	n	$2n$
B_n	$\frac{(-1)^n}{2^n n!}$	$\frac{1}{n!}$	$(-1)^n$
a_n	$\frac{\Gamma(\alpha + \beta + 2n + 1)}{2^n n! \Gamma(\alpha + \beta + n + 1)}$	$\frac{(-1)^n}{n!}$	2^n
b_n	$\frac{(\alpha - \beta)\Gamma(\alpha + \beta + 2n)}{2^n (n-1)! \Gamma(\alpha + \beta + n + 1)}$	$(-1)^{n-1} \frac{\alpha + n}{(n-1)!}$	0
d^2	$\frac{2^{\alpha+\beta+1} \Gamma(\alpha + n + 1) \Gamma(\beta + n + 1)}{n! (\alpha + \beta + 2n + 1) \Gamma(\alpha + \beta + n + 1)}$	$\frac{\Gamma(\alpha + n + 1)}{n!}$	$2^n n! \sqrt{\pi}$
α_n	$\frac{2(n+1)(\alpha + \beta + n + 1)}{(\alpha + \beta + 2n + 1)(\alpha + \beta + 2n + 2)}$	$-(n+1)$	$\frac{1}{2}$
β_n	$\frac{\beta^2 - \alpha^2}{(\alpha + \beta + 2n)(\alpha + \beta + 2n + 2)}$	$\alpha + 2n + 1$	0
γ_n	$\frac{2(\alpha + n)(\beta + n)}{(\alpha + \beta + 2n)(\alpha + \beta + 2n + 1)}$	$-(\alpha + n)$	n

C. Mathematica File

We have discussed basic potentials of the nonrelativistic and relativistic quantum mechanics that can be integrated in the Nikiforov and Uvarov paradigm with the aid of the Mathematica computer algebra system. (The corresponding Mathematica notebook is available from the authors by a request. It is also posted on Wolfram community <https://community.wolfram.com/groups/-/m/t/2897057> and featured in the editorial columns <https://community.wolfram.com/content?curTag=staff+picks> .)

In section 2, the general formulas are derived. In the notebook, they are stored in global variables that will be used in all the subsequent sections. For this purpose, allow Mathematica to evaluate all initialization cells. After that one can run each case independently from the others. The results, for the most integrable cases that are available in the literature, are presented in the Tables 1–18.

The Mathematica package, `HolonomicFunctions`, can be downloaded from the webpage of the third-named author: <http://www.koutschan.de/>.

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